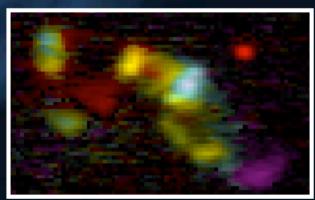
# Are There Dark Galaxies in the Local Group?

Josh Simon Caltech

Marla Geha (HIA) Tim Robishaw (Berkeley) Leo Blitz (Berkeley)







### **DARKGALAXY**

Dark Galaxy is a massively multiplayer browser based strategy game.

#### **Welcome to Dark Galaxy**

In Dark Galaxy you control planets and fleets in games lasting for 3 months at a time. Read more..

Click here to join the game!

Jupiter Login (game open)

Visit our forums | Game manual



## Outline

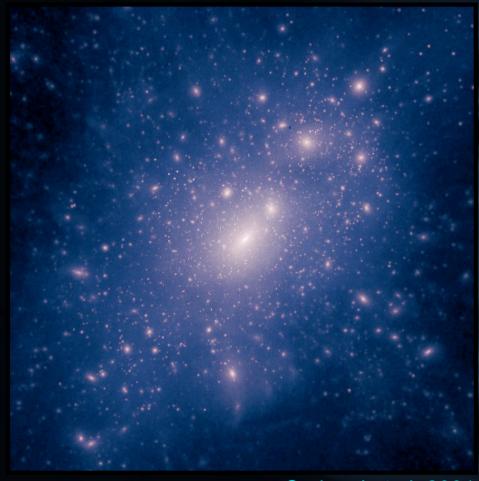
- I. The missing satellite problem
- II. Can the new SDSS dwarfs be the missing satellites?

Keck/DEIMOS kinematics of 8 SDSS dSphs

III. Are HVCs the missing satellites?

Arecibo H I maps of 12 HVCs

# CDM and the Missing Satellites



Springel et al. 2001

CDM predicts large numbers of subhalos (~100-1000 for a Milky Way-sized galaxy)

Milky Way only has 23 known satellites

What happened to the rest of them?

# CDM and the Missing Satellites



CDM predicts large numbers of subhalos (~100-1000 for a Milky Way-sized galaxy)

Milky Way only has 23 known satellites

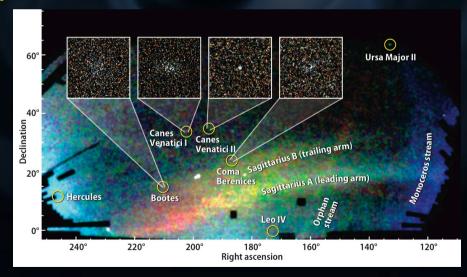
What happened to the rest of them?

Springel et al. 2001

## **New SDSS Dwarfs**

- Only 11 Milky Way dwarfs known through 2004
- Since 2005, SDSS has discovered:
  - **8 new dSphs** Willman et al. (2005), Zucker et al. (2006a,b), Belokurov et al. (2006,2007)
  - **1 dirr** Irwin et al. (2007)
  - 3 ??? dSph/GCs?

Willman et al. (2005), Belokurov et al. (2007), Walsh et al. (2007)



Belokurov et al. (2006)

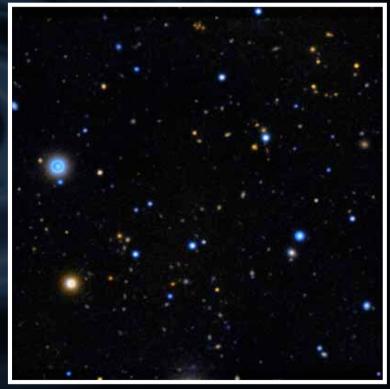
## **New SDSS Dwarfs**

### Old dwarfs (Leo II)



Palomar Sky Survey

### New dwarfs (Ursa Major I)



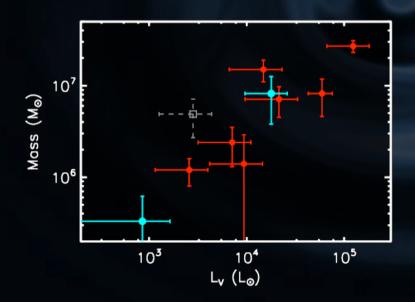
Willman et al. (2005)

# Keck Spectroscopic Survey of SDSS Dwarfs

Medium-resolution spectra of 841 stars (424 members) across 8 dwarfs

(CVn I, CVn II, Coma Berenices, Hercules, Leo IV, Leo T, UMa I, UMa II)

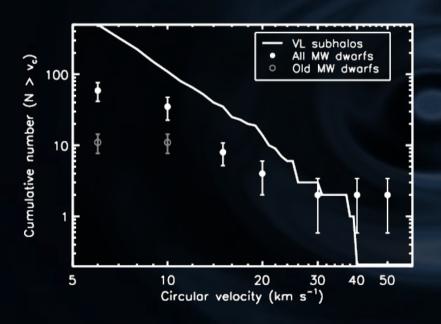
Measured stellar velocities + metallicities

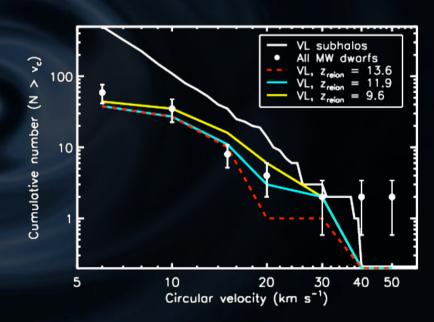


- Simon & Geha (2007)
- □ UMa II (tidally disrupting; SG07)
- Martin et al. (2007)

# Are There Still Satellites Missing?

Comparison to Via Lactea N-body simulation (Diemand et al. 2007)





Simon & Geha (2007)

## What About HVCs?

### Major HVC models:

Extra/circumgalactic	Hot halo	Tidal debris
(Oort 1966, Blitz et al. 1999)	(Maller & Bullock 2004)	(no formal model)
- d = 50 - 250 kpc	- d ~ 150 kpc	- d = 10 - 50 kpc
$- m = 10^5 - 10^7 M_{\odot}$	$- m = 10^5 - 10^7 M_{\odot}$	$- m = 10^3 - 10^5? M_{\odot}$
- dark matter halos	- no dark matter	- no dark matter
- gravitationally bound	- pressure-confined	- not bound

### Crucial questions about the bulk of the HVC population:

- Are they gravitationally bound?
- Do they have dark matter halos?

### Are HVCs Bound?

If not, calculate expansion timescale

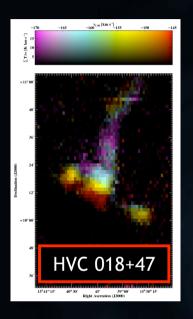
$$\Delta t \approx \frac{2r}{\Delta v} \left( \frac{d}{100 \ kpc} \right)$$

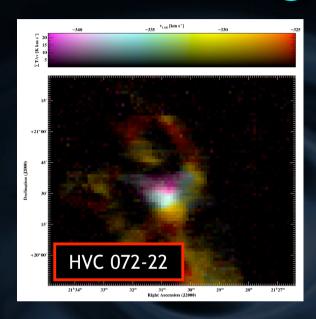
- $5.3 \times 10^7 \text{ yr} < \Delta t < 2.5 \times 10^8 \text{ yr}$
- Not stable for very long
- What about dynamical masses?
  - Assuming virialization:

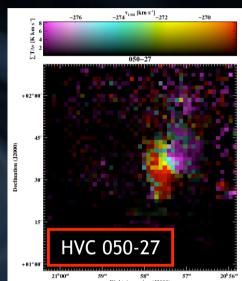
300 (d/100 kpc)<sup>-1</sup> <  $M_{vir}/M_{HI}$  < 2100 (d/100 kpc)<sup>-1</sup>

Must have dark matter to be bound!

# Four Rotating HVCs?

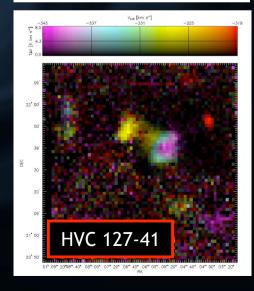






Simon et al. (in prep)

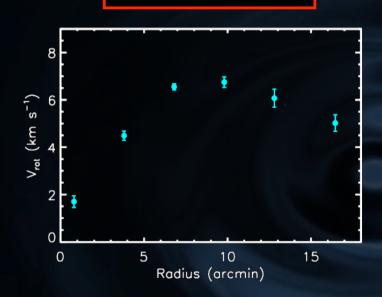
4/12 HVCs have major axis velocity gradients with apparent rotation velocities of 4 - 14 km/s . . .



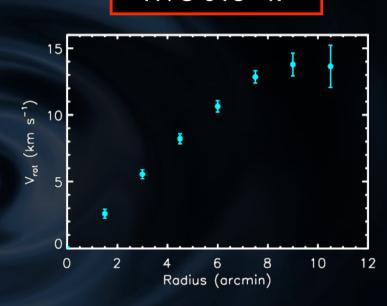
Robishaw, Simon & Blitz (2002)

## **HVC Rotation Curves**

HVC 127-41



 $V_{\text{rot}} = 7 \text{ km s}^{-1}$  $M_{\text{dyn}}/M_{\text{HI}} = 7 (d/700 \text{ kpc})^{-1}$  HVC 018+47



$$V_{\text{rot}} = 14 \text{ km s}^{-1}$$
  
 $M_{\text{dyn}}/M_{\text{HI}} = 78 \text{ (d/100 kpc)}^{-1}$ 

## Summary

- 1) Ultra-faint dwarfs are extremely dark matter-dominated
  - Masses of  $10^6 10^7 M_{\odot}$ , M/L =  $100 1000 M_{\odot}/L_{\odot}$
  - Ultra-faint dwarfs alleviate but don't remove the substructure problem . . .
  - UNLESS dwarf galaxy formation was strongly suppressed by reionization at  $z \approx 12$  (Simon & Geha 2007)
- 2) HVC kinematics show that they are either dark matter-dominated or unbound
  - Some HVCs appear to be plausible dark galaxies

Solution to missing satellite problem now in reach . . .