

Physics and chemistry of star formation in the outer Galaxy

PhD student: Diego Gigli

Supervisor: Francesco Fontani

A little bit of me

who: Diego Gigli

where: Ronta (Mugello valley)

when: March 1998



2021: Bachelor's degree in Physics and Astrophysics

2024: Master's degree in Physical and Astrophysical Sciences

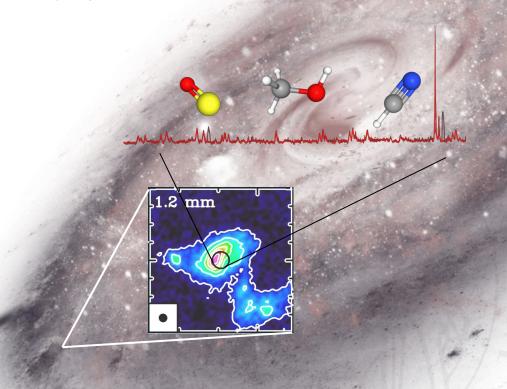
????: PhD in Physics and Astronomy

The project CHEMical complexity of star-forming regions in the OUTer galaxy (CHEMOUT)

Scientific goal

Understanding how chemistry changes throughout the Milky Way and revealing the chemical composition of star-forming regions in the outer Galaxy

Fontani+2022a; Fontani+2022b; Colzi+2023; Fontani+2024; Gigli+2025



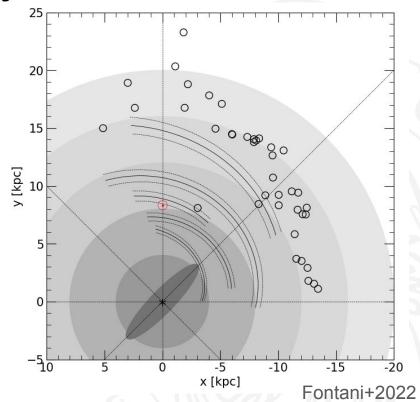
The CHEMOUT project

Observation sample

35 star-forming regions in the outer Galaxy between 9 and 24 kpc from the Galactic center

Observational facilities

IRAM 30m, Yebes 40m, GBT, ALMA, and NOEMA (so far...)



CHEMOUT: results

Fontani+2022,2022b. CHEMOUT I, and II Abundances of key organic species does not vary significantly with Galactocentric radius

Colzi+2023. CHEMOUT III

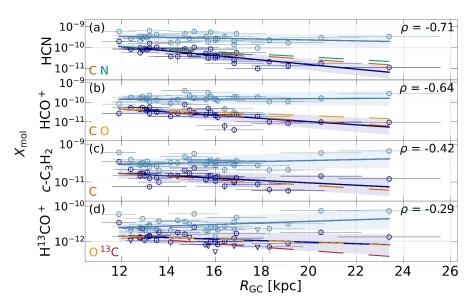
¹⁴N/¹⁵N ratio supports the hypothesis that ¹⁵N isotopes comes from novae explosions

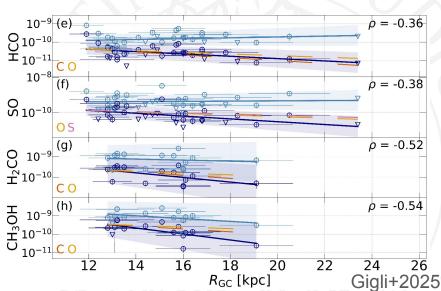
Fontani+2024. CHEMOUT IV

The outermost region (WB89-670 @ 23.4 kpc) shows local chemical differentiation and an unexpectedly high Carbon [C/H] abundance

CHEMOUT V. Chemical composition gradients as a function of the Galactocentric radius

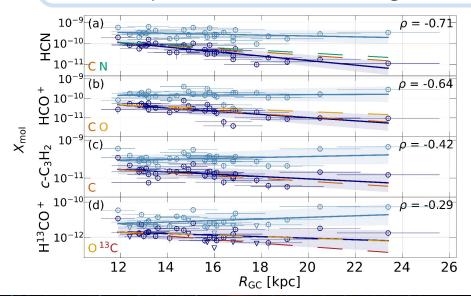
We estimated how the abundances of some species changes through the Galactic disk, comparing them with the elemental abundances

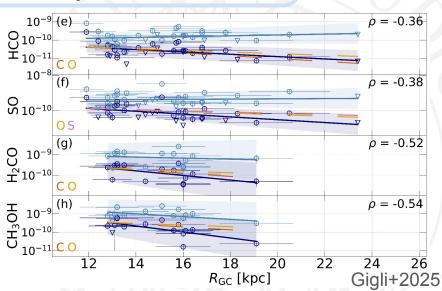




CHEMOUT V. Chemical composition gradients as a function of the Galactocentric radius

The abundances of most of the species scales at most as the elemental abundance of their parent element. The formation efficiency of these species is constant throughout the Galaxy.







The PhD project

Yebes 40m observations

Year 0 abundances from emission lines

iram

iram

Year 1
kinetic temperatures from ammonia
emission

Year 2
density profiles from dust emission
maps



high-resolution observations
with NOEMA



Year 3
observations - model
comparison with UCL chem

new observations

What's next?

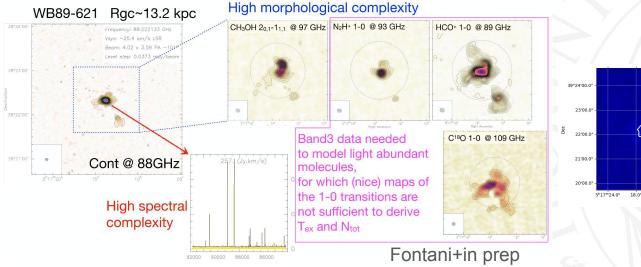
Spaccino+in prep → Estimating the gas-to-dust ratio in the outer Galaxy

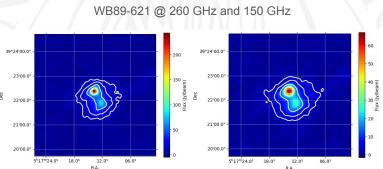
Gigli+in prep → Ammonia analysis for the gas temperature

Gigli+in prep → Morphology study from the continuum maps

Fontani+in prep → MORGAN project: NOEMA high-resolution analysis of WB89-621

Colzi+in prep → ¹³C/¹²C ratio in the outer Galaxy





Gigli+in prep

A broader impact

- Redefine the concept of Galactic Habitable Zone
- Identify prime targets for next-generation radio/mm telescopes (ngVLA, SKA)
- Study low-metallicity environments as the early Milky Way





Thank you for your attention Physics and chemistry of star formation in the outer Galaxy

Diego Gigli (diegogigli.dg@gmail.com)