

The baryon cycle: accretion, outflow and the circum-galactic medium

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*Celine Peroux*

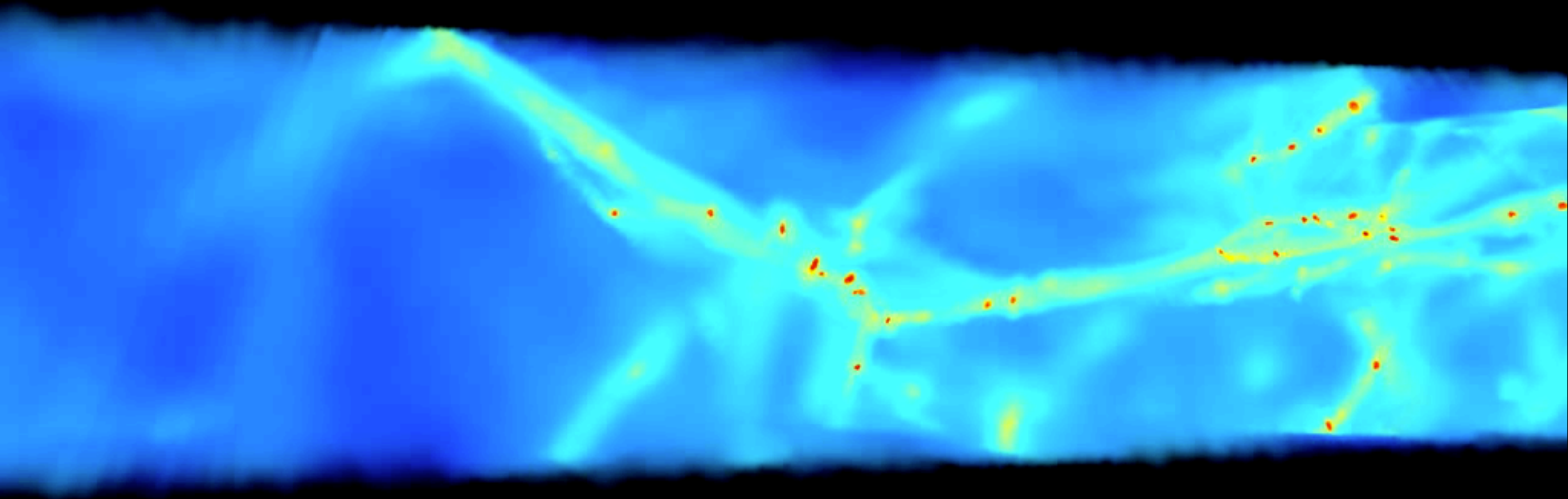
*Laboratoire d'Astrophysique de Marseille, France*

# PLAN

1. Baryons traced by Neutral Gas
2. Gas Flows in Absorption
3. Conclusion



*Stephan Frank*



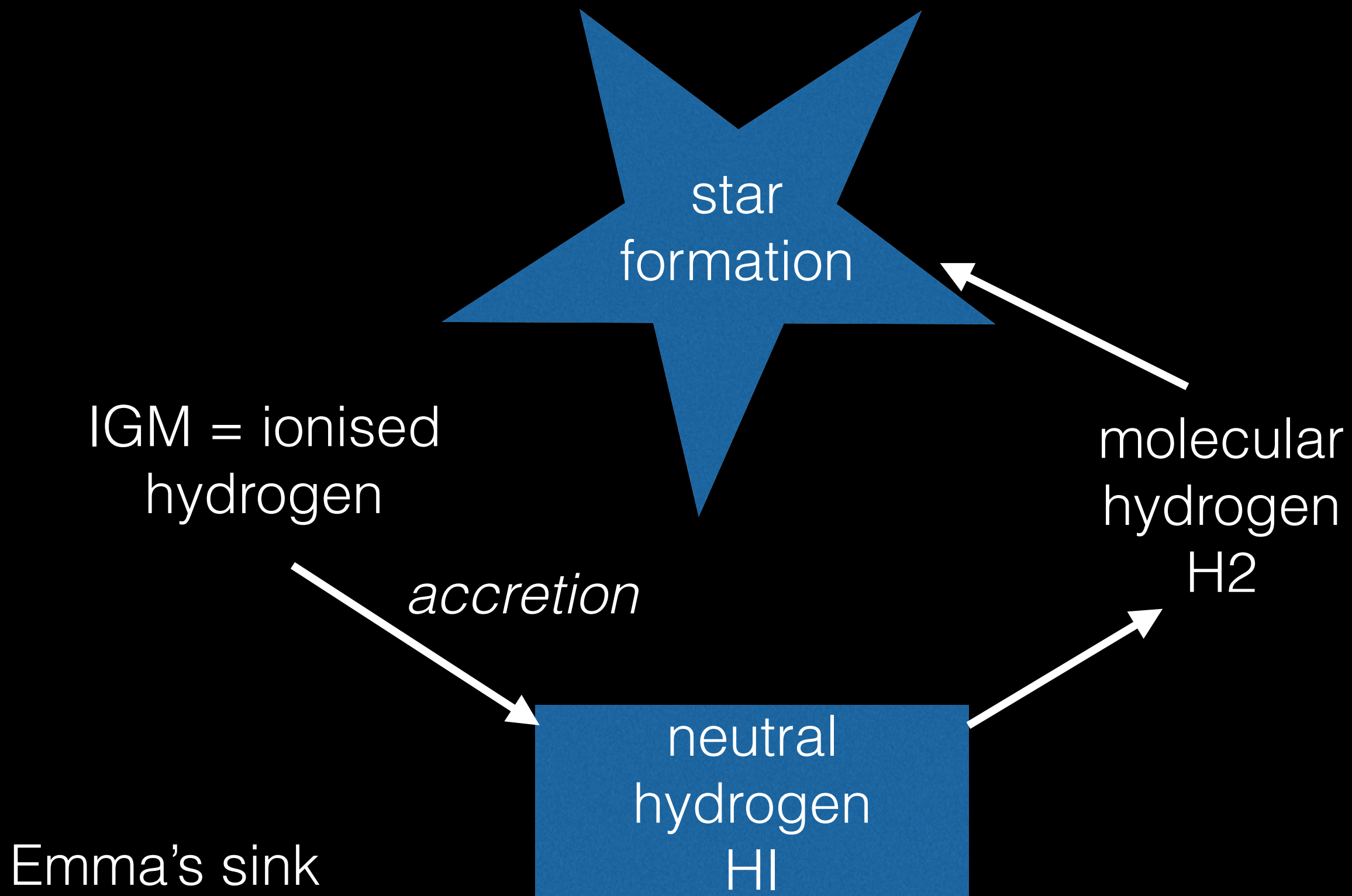
RAMSES AMR simulations (Teyssier+)  
Ly $\alpha$ ,  $z=0.75$ ,  $6.6 \times 6.6 \times 300$  (phys) Mpc

Frank..CP+12

# PLAN

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# Why Neutral Gas HI?





*Attila Popping*

# The Data

- ESO UVES Advanced Data Products
- 250 quasar spectra, 1500 hrs of VLT time
- 150 DLAs/sub-DLAs

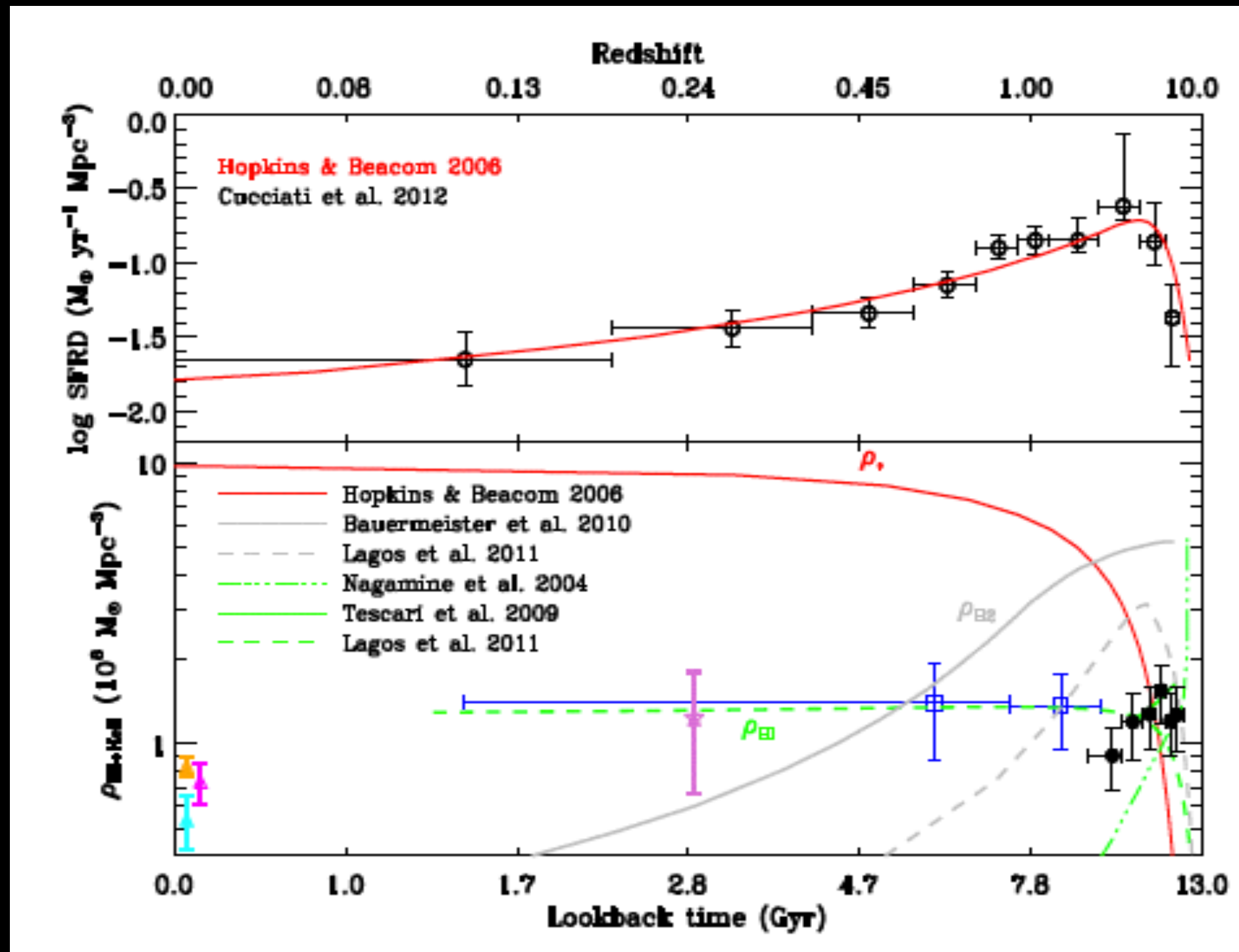
Zafar, Popping, CP+13a



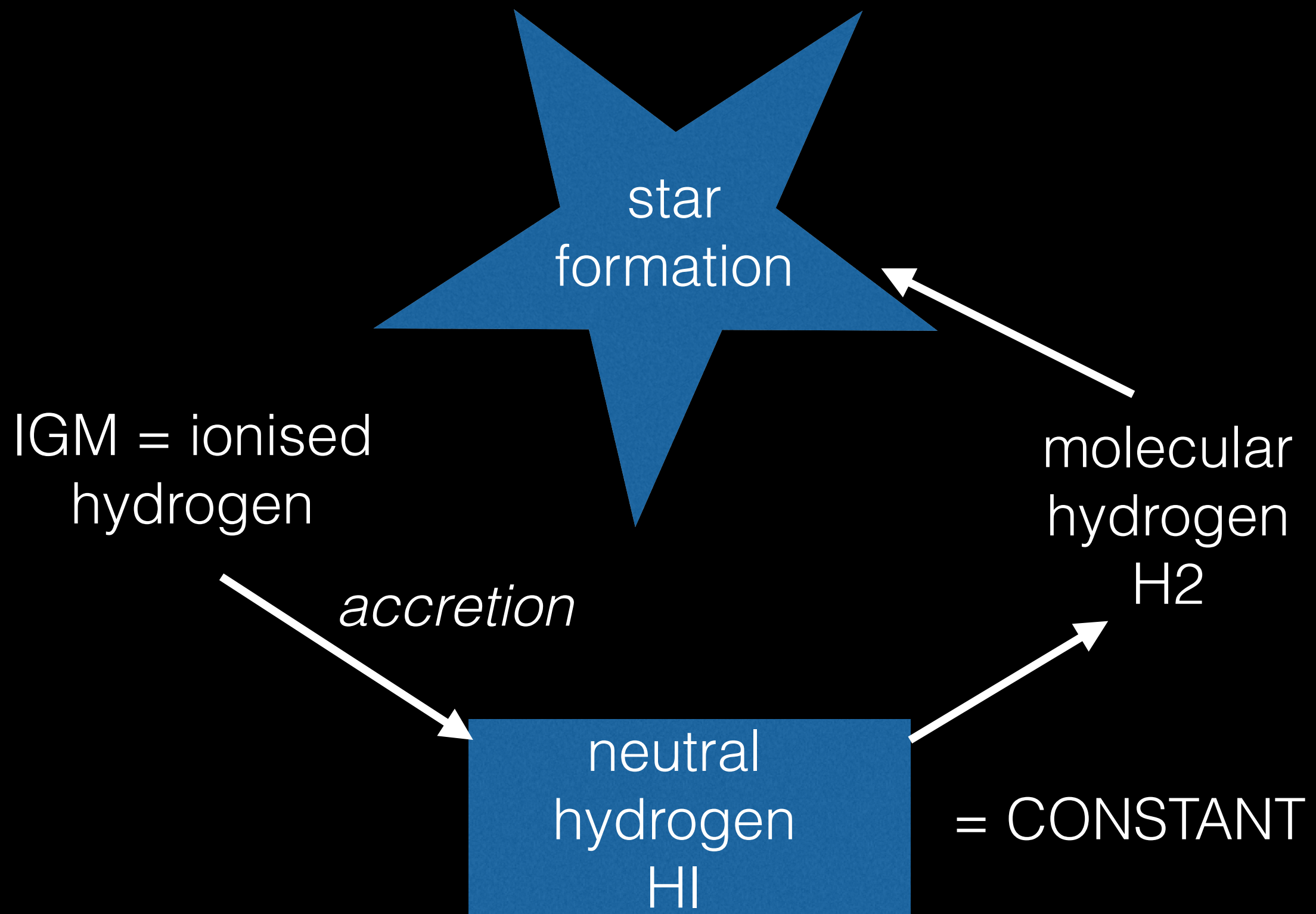
# Neutral Hydrogen Gas Mass

Tayyaba Zafar

- 10% of gas forming stars
- cosmic coincidence

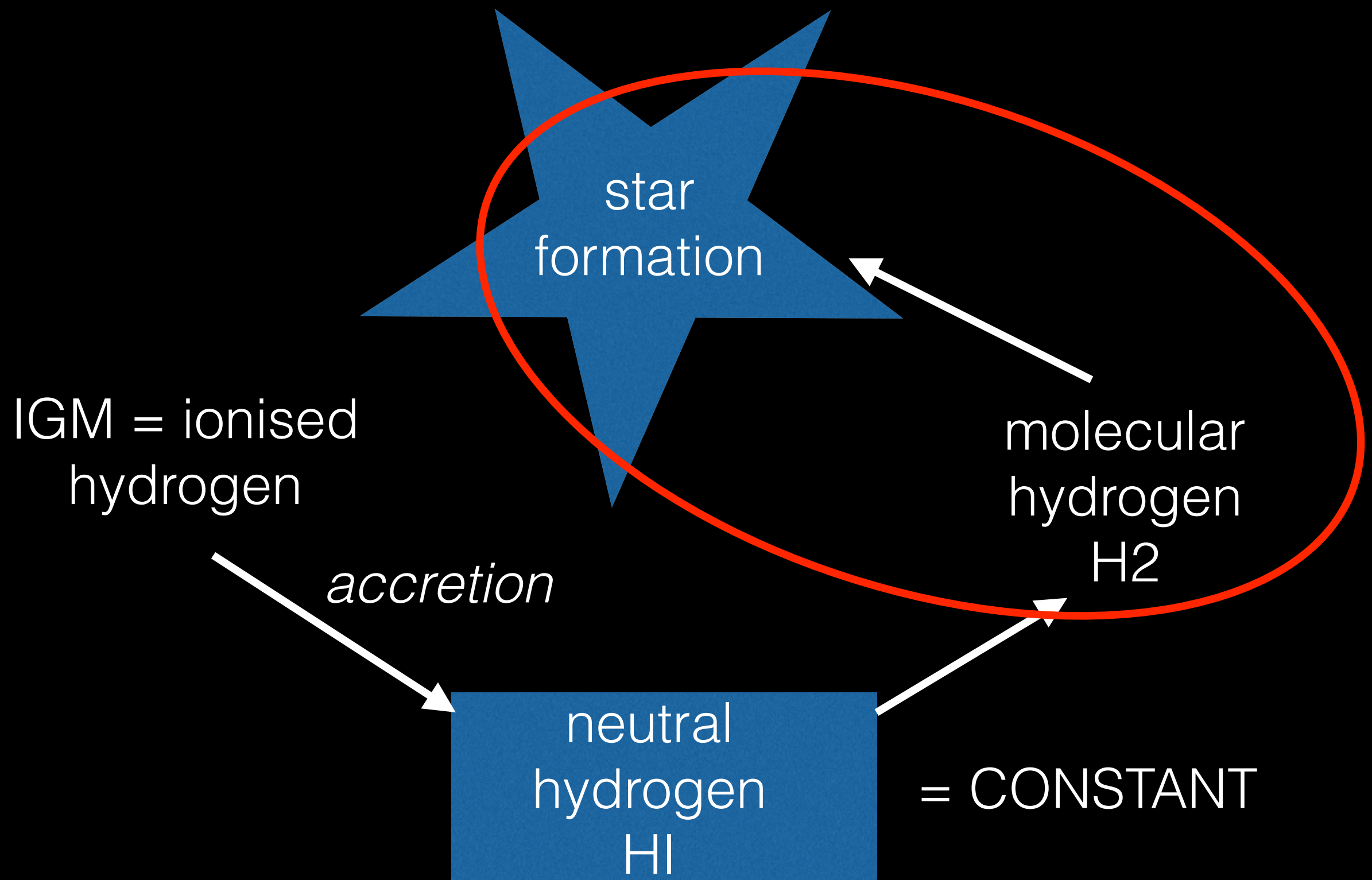


# Neutral Gas HI

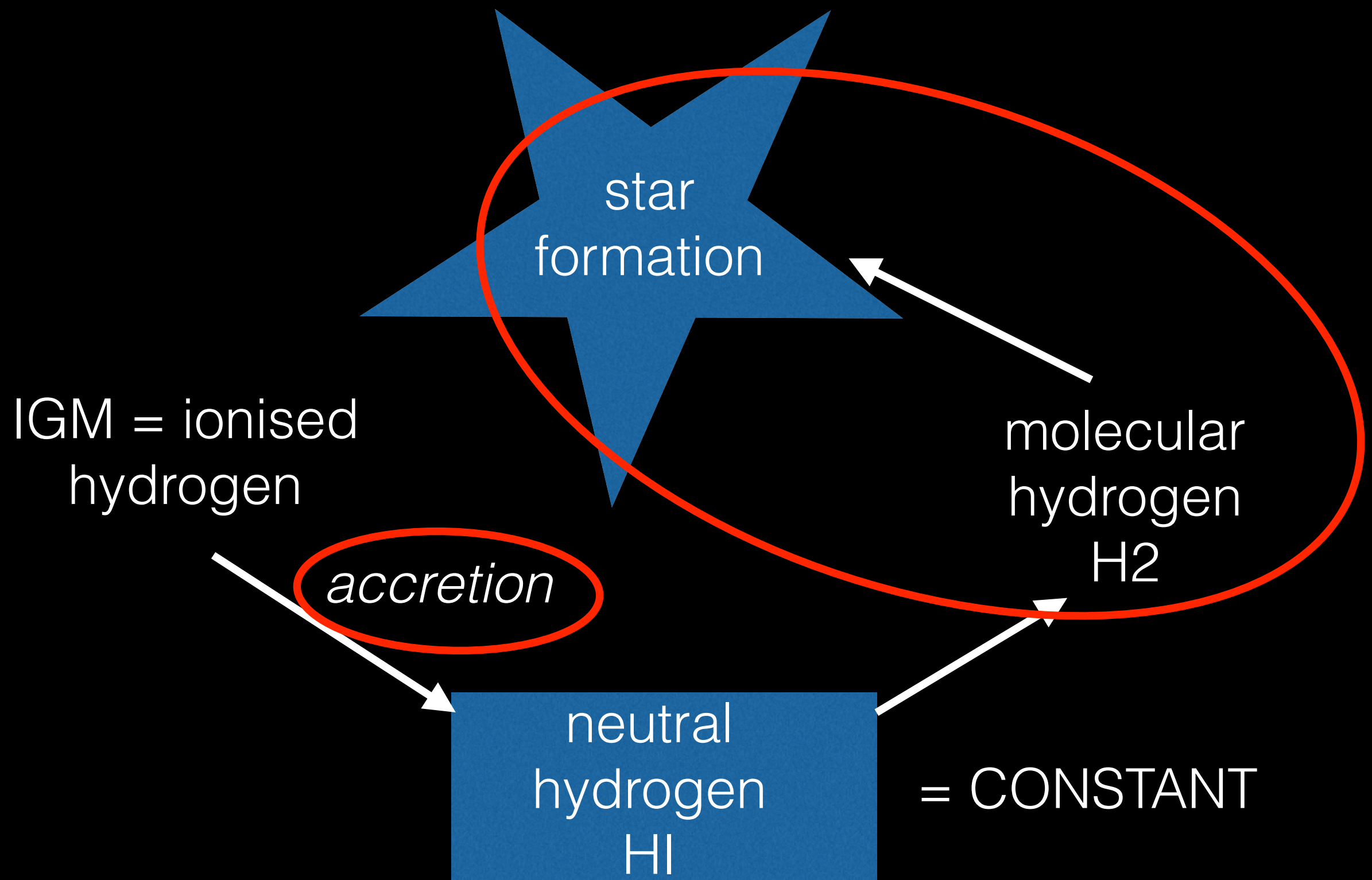




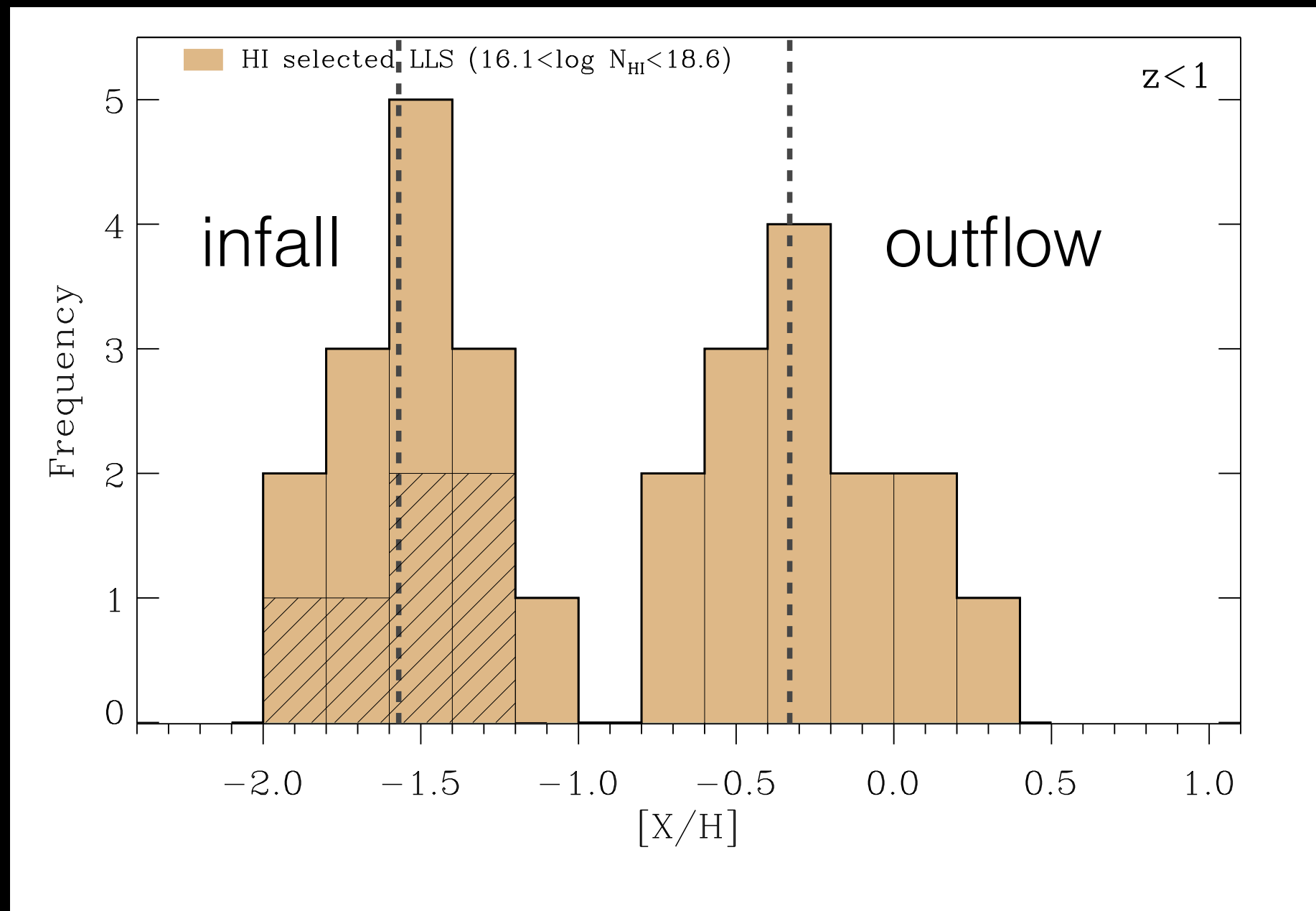
# Neutral Gas HI



# Neutral Gas HI



# Bimodal Metallicity



LLS at  $z \sim 1$

Lehner+13, Wotta+15

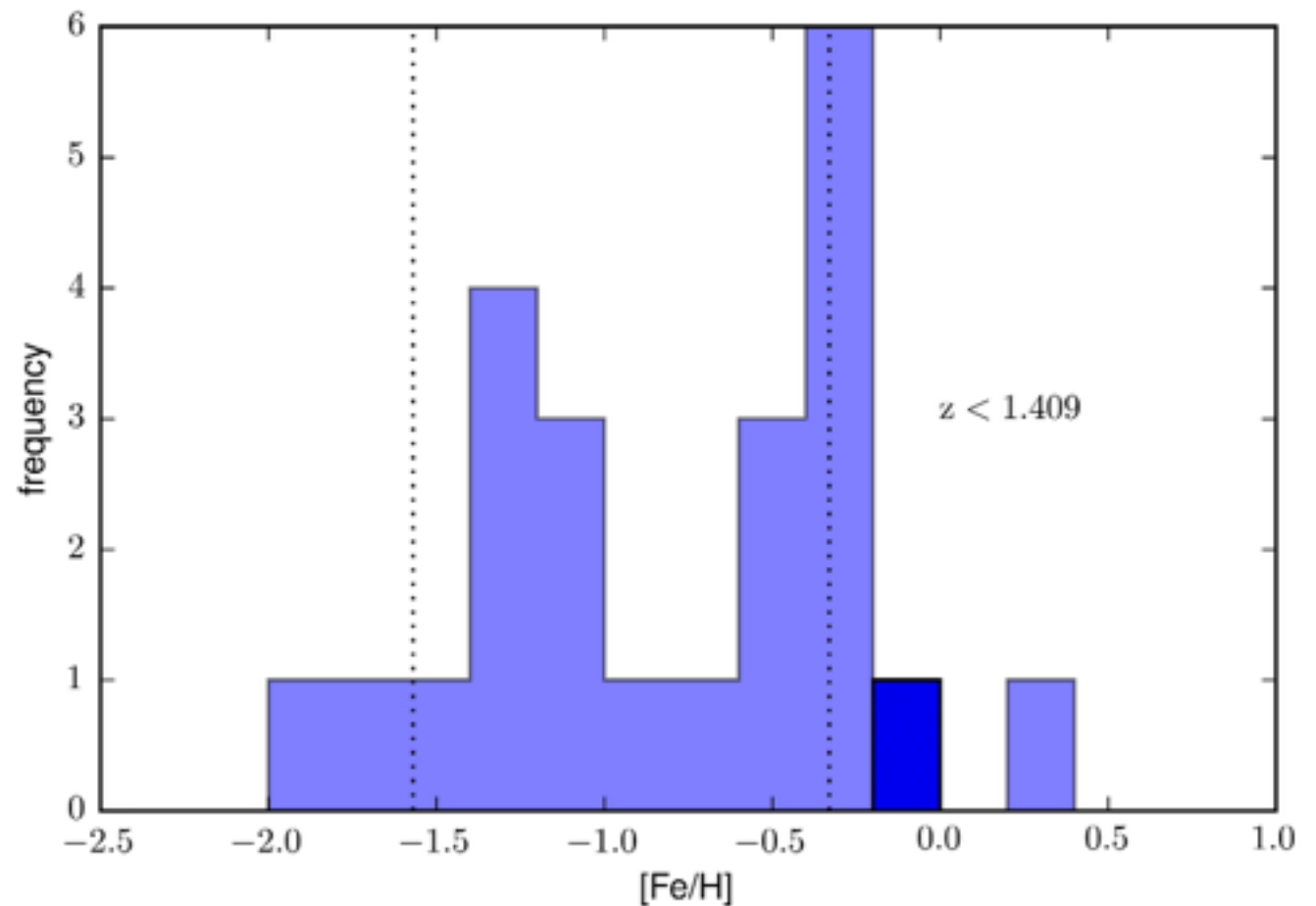


Samuel Quiret

# Metallicity Distribution

- hint of bimodality?

See Samuel's poster



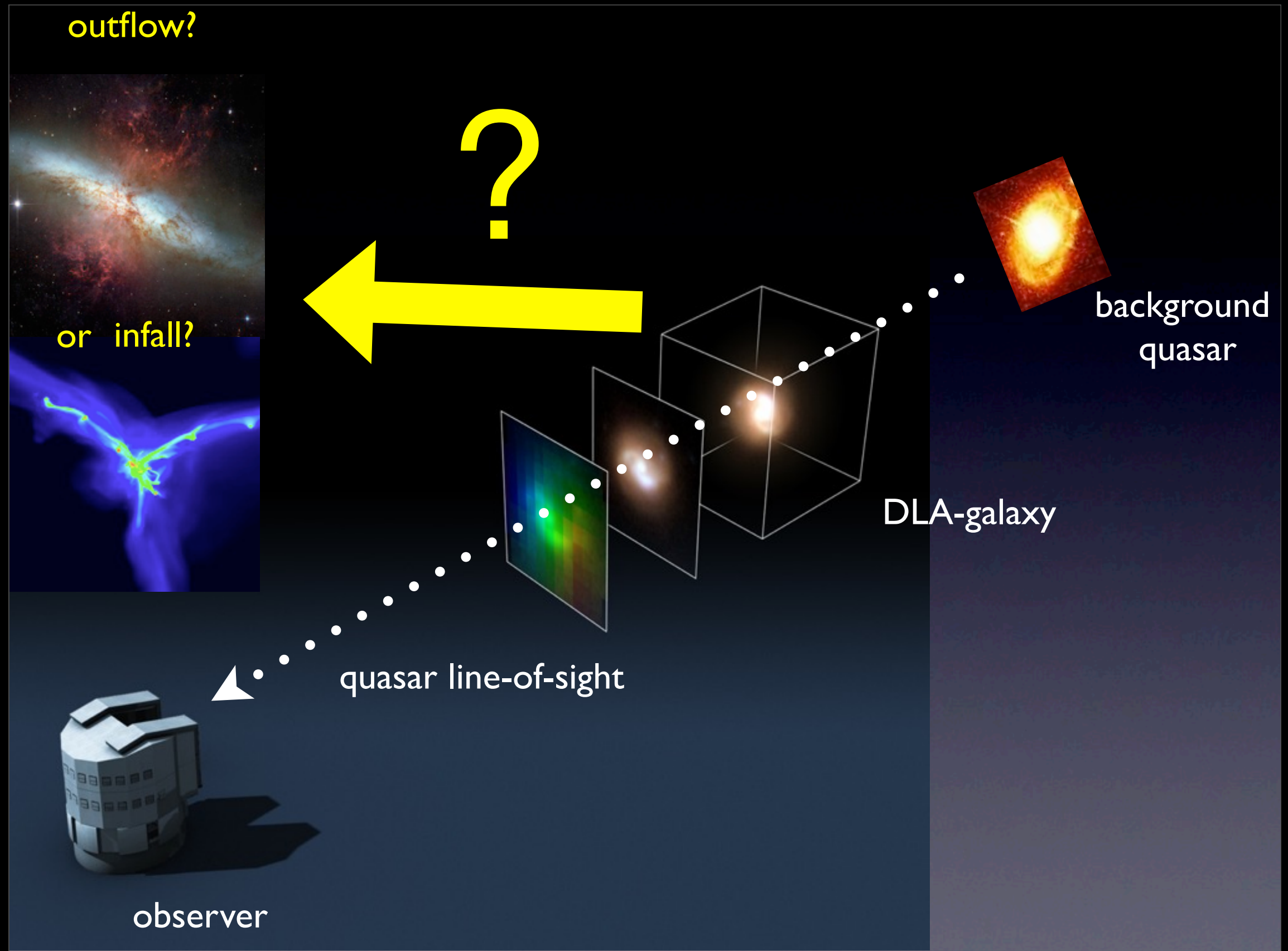
sub-DLAs at  $z < 1.4$

Quiret, CP+15

# PLAN

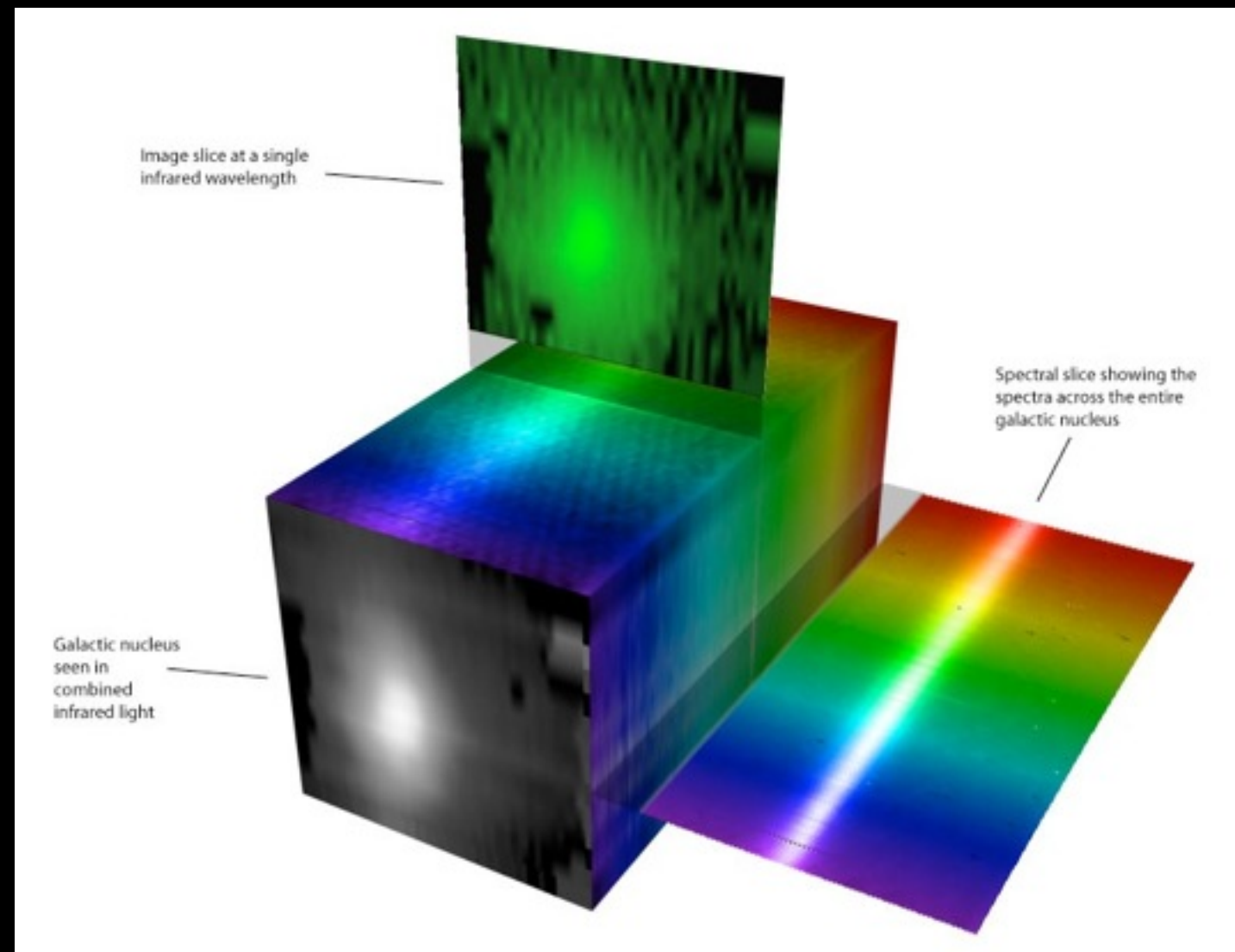
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# Probing Gas Flows



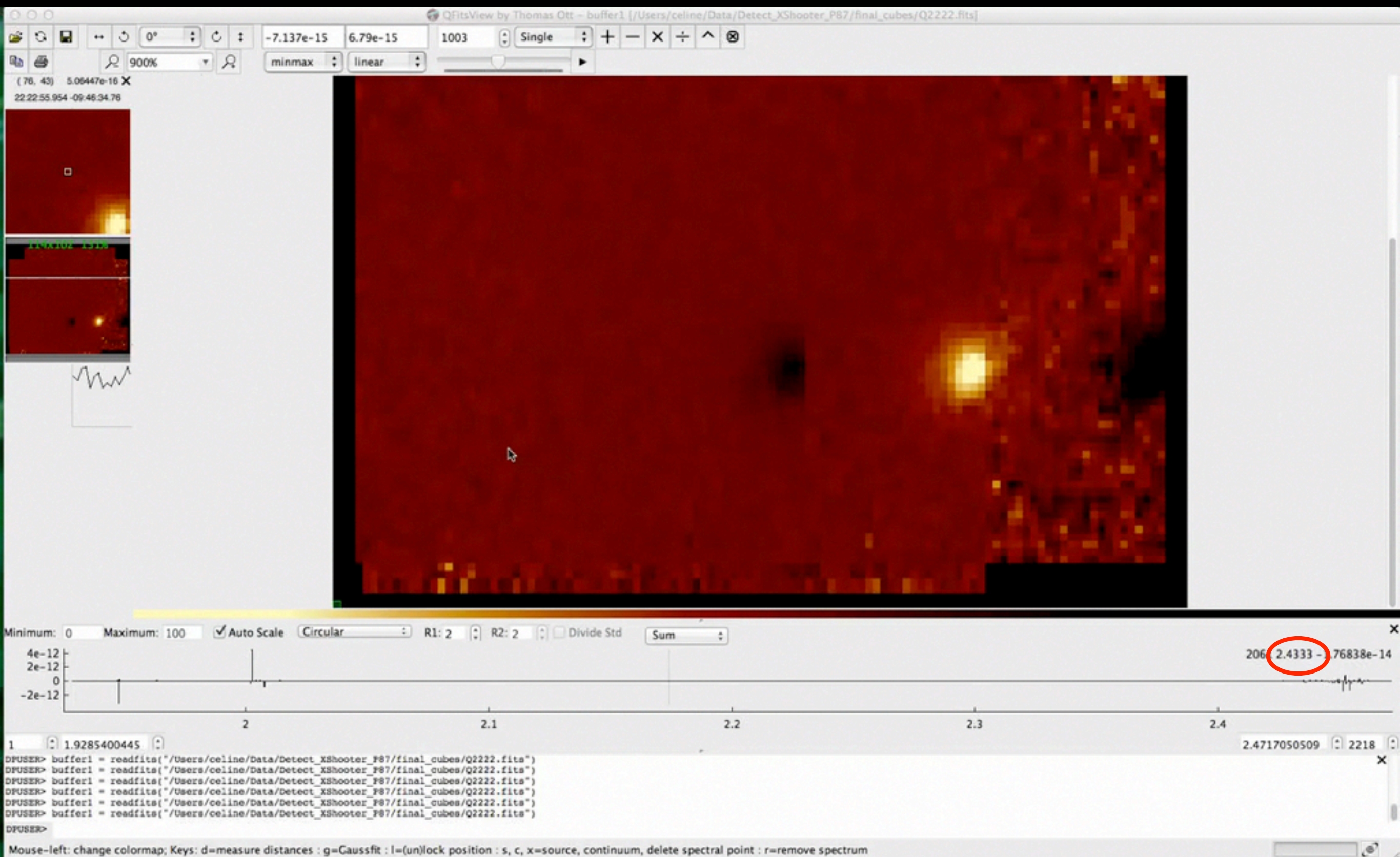
# Observational Strategy: Kinematics

- search for stellar content of absorbers with known N(HI)
- IFU allows to remove signature of background quasar



VLT/SINFONI





- spectral PSF subtraction (SPSF)

CP+12

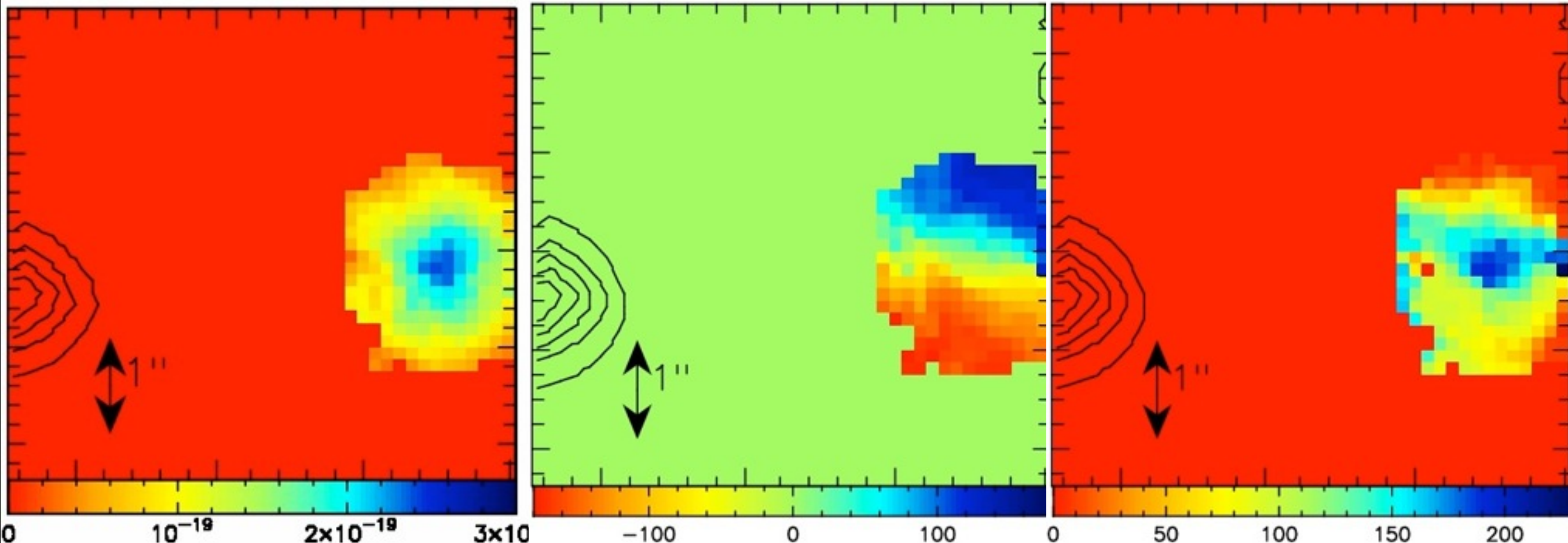


# Kinematics

Q1009 H $\alpha$ (z=0.887)

Q1009 Vel(km/s)

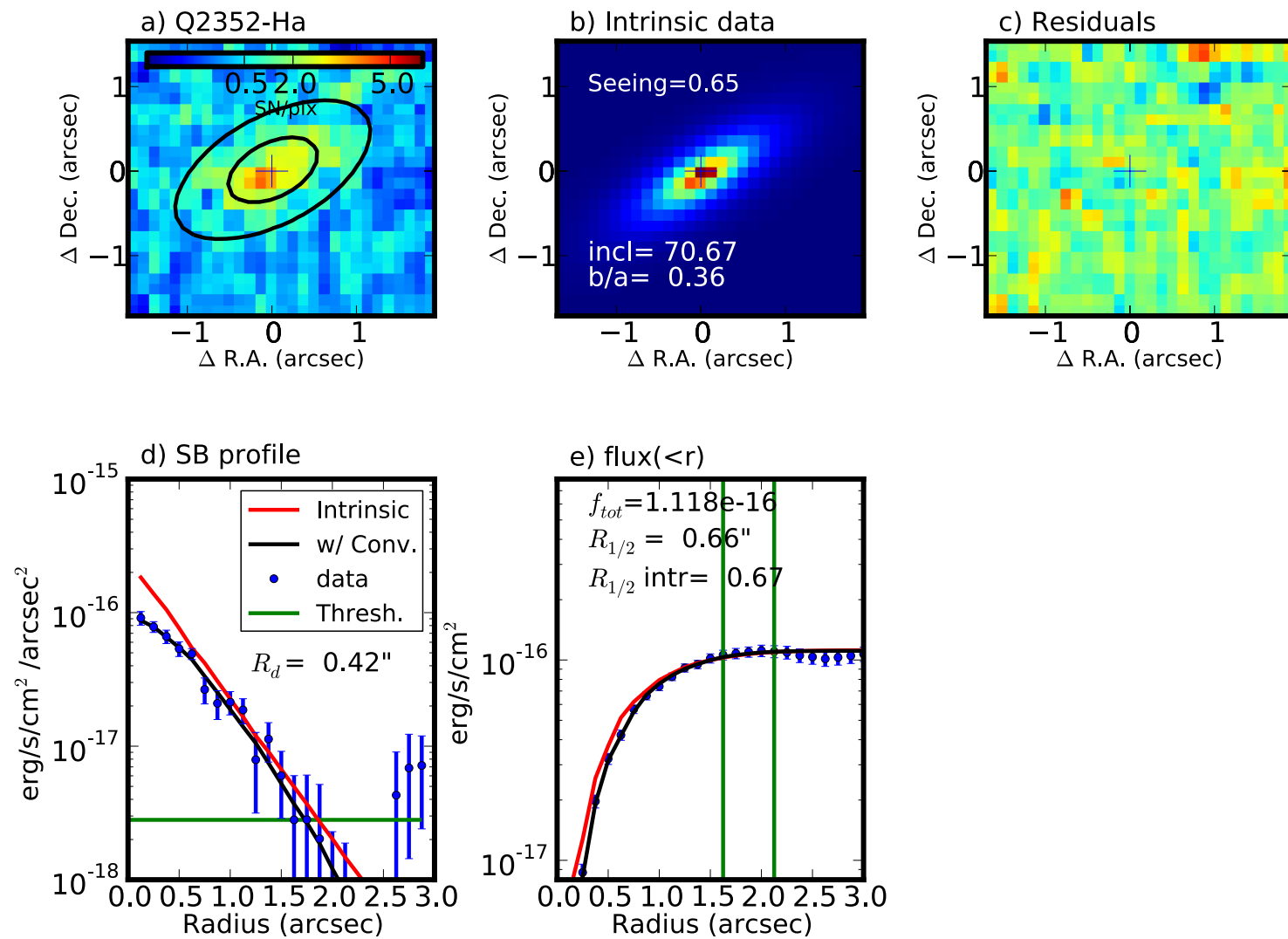
Q1009 Disp(km/s)



- orientation, inclination
- velocity, dispersion

CP+11b

# Mass Estimates



# Previous Results

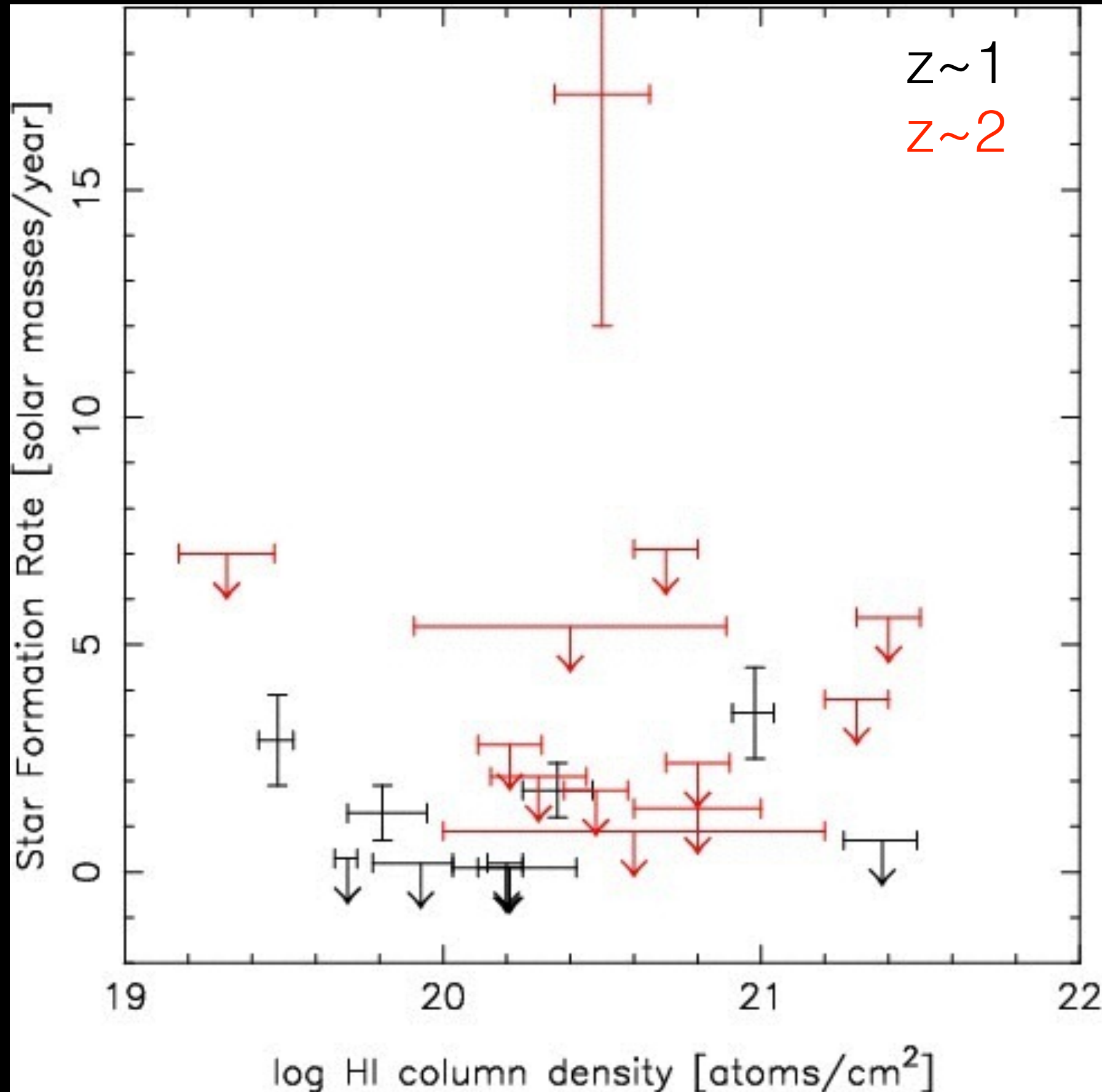
| Quasar     | Galaxy Orientation | b [kpc] | Direction to quasar line-of-sight aligned with | $V_{\max}$ [km/s] | $\Delta v$ [km/s] | Absorption Profile              | Conclusion                         |
|------------|--------------------|---------|------------------------------------------------|-------------------|-------------------|---------------------------------|------------------------------------|
| Q0302–223  | edge-on            | 25      | minor axis                                     | 11                | 120               | doubled-peaked                  | $\Rightarrow$ co-rotating/outflow? |
| Q0452–1640 | face-on?           | 16      | major axis                                     | 100               | 230               | either-side of $z_{\text{gal}}$ | $\Rightarrow$ merger/outflow?      |
| Q1009–0026 | edge-on            | 39      | minor axis?                                    | 250               | 334               | asymmetrical                    | $\Rightarrow$ outflow              |
| Q2222–0946 | edge-on            | 6       | n/a <sup>†</sup>                               | 20                | 200               | centred and complex             | $\Rightarrow$ outflow              |
| Q2352–0028 | edge-on            | 12      | major axis                                     | 140               | 220               | centred and complex             | $\Rightarrow$ co-rotating/outflow? |

<sup>†</sup>: in the case of Q2222–0946, the major axis is undefined because of the compact nature of the galaxy.

$\Rightarrow$  in 2 cases, we have strong indications of outflows

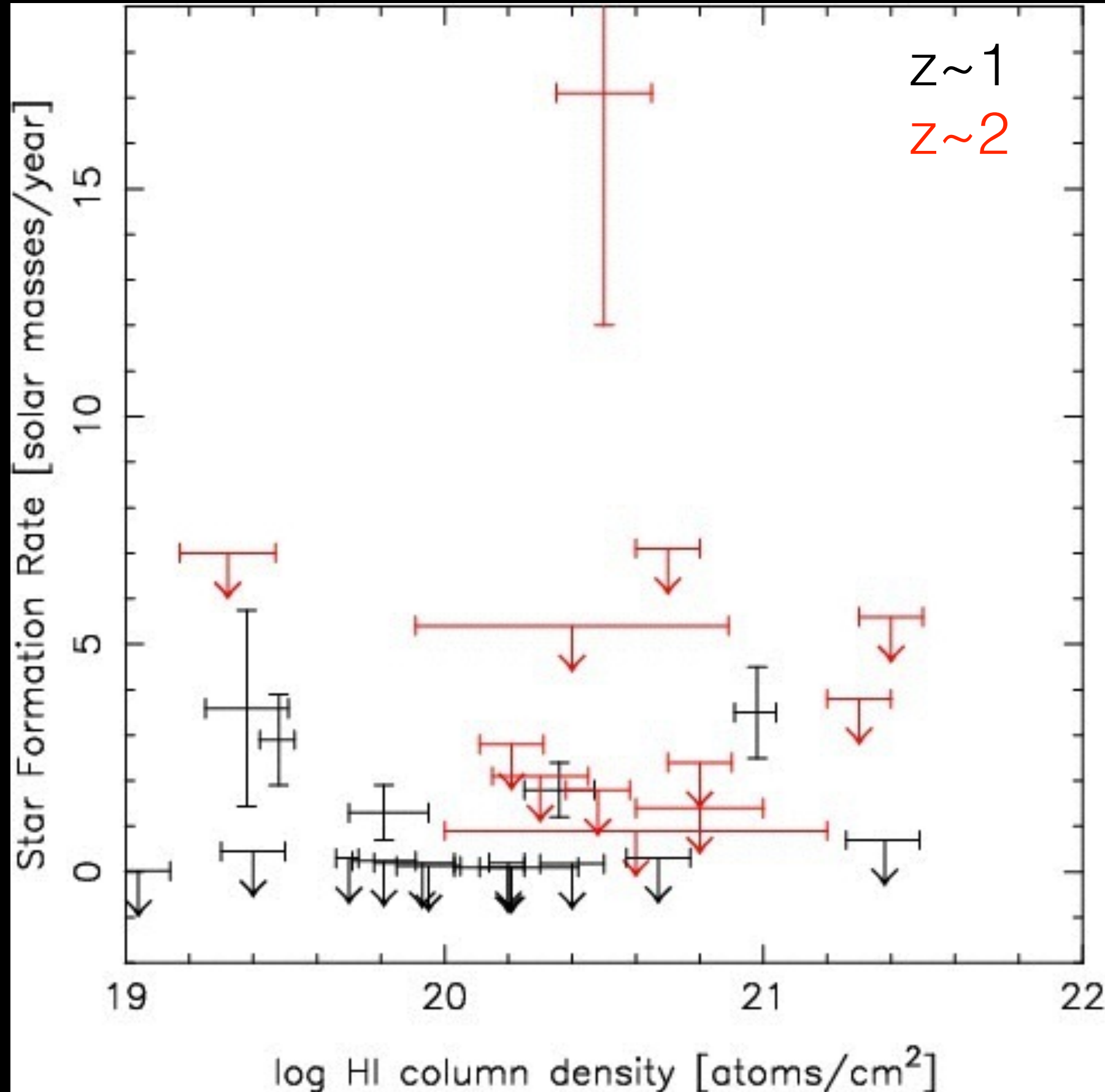
# Star Formation Rates

- SFR  $\sim$  few  $M_{\odot}/\text{yr}$
- $b \sim 10\text{-}40$  kpc



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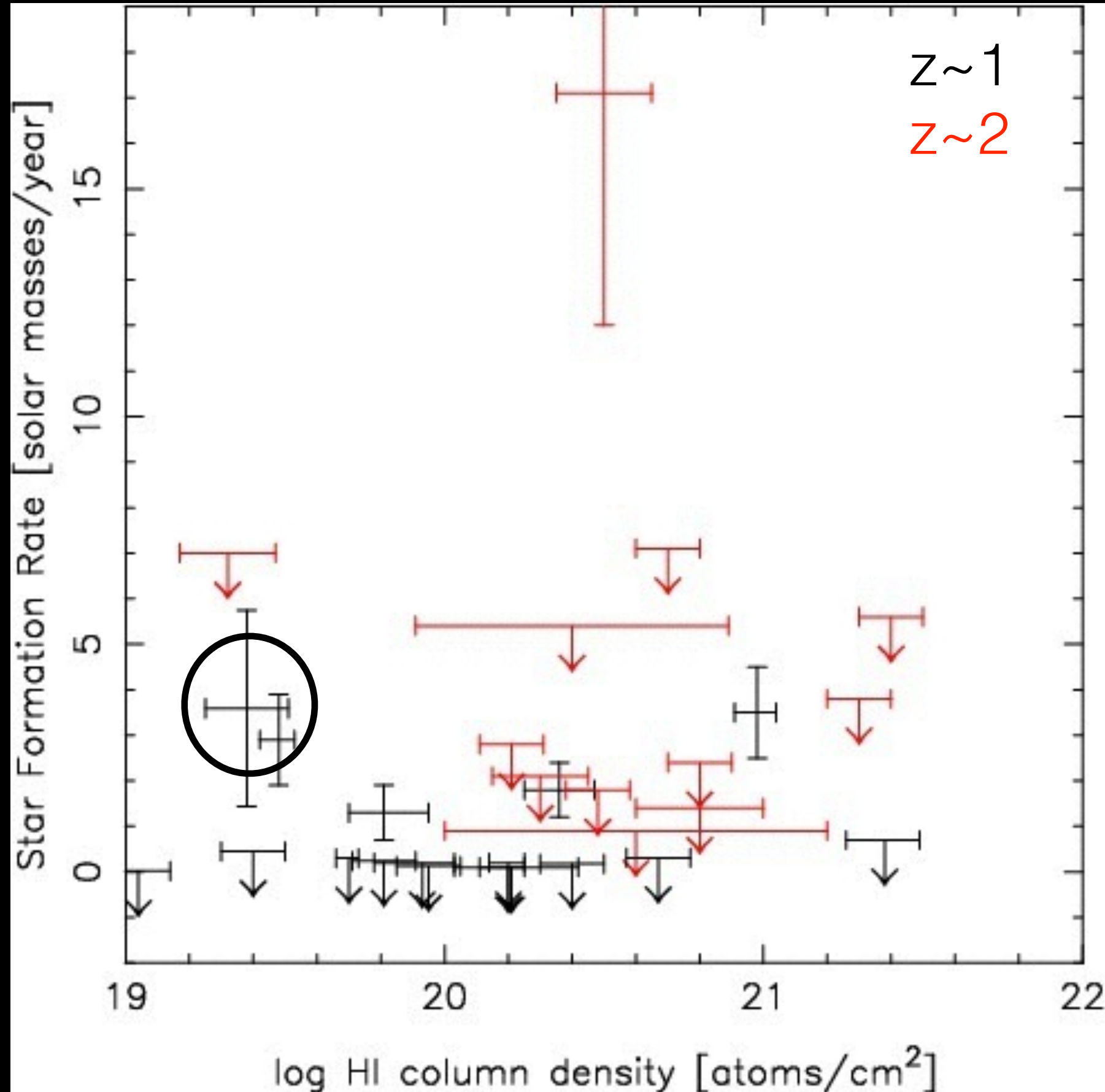
CP+ in prep



# Star Formation Rates

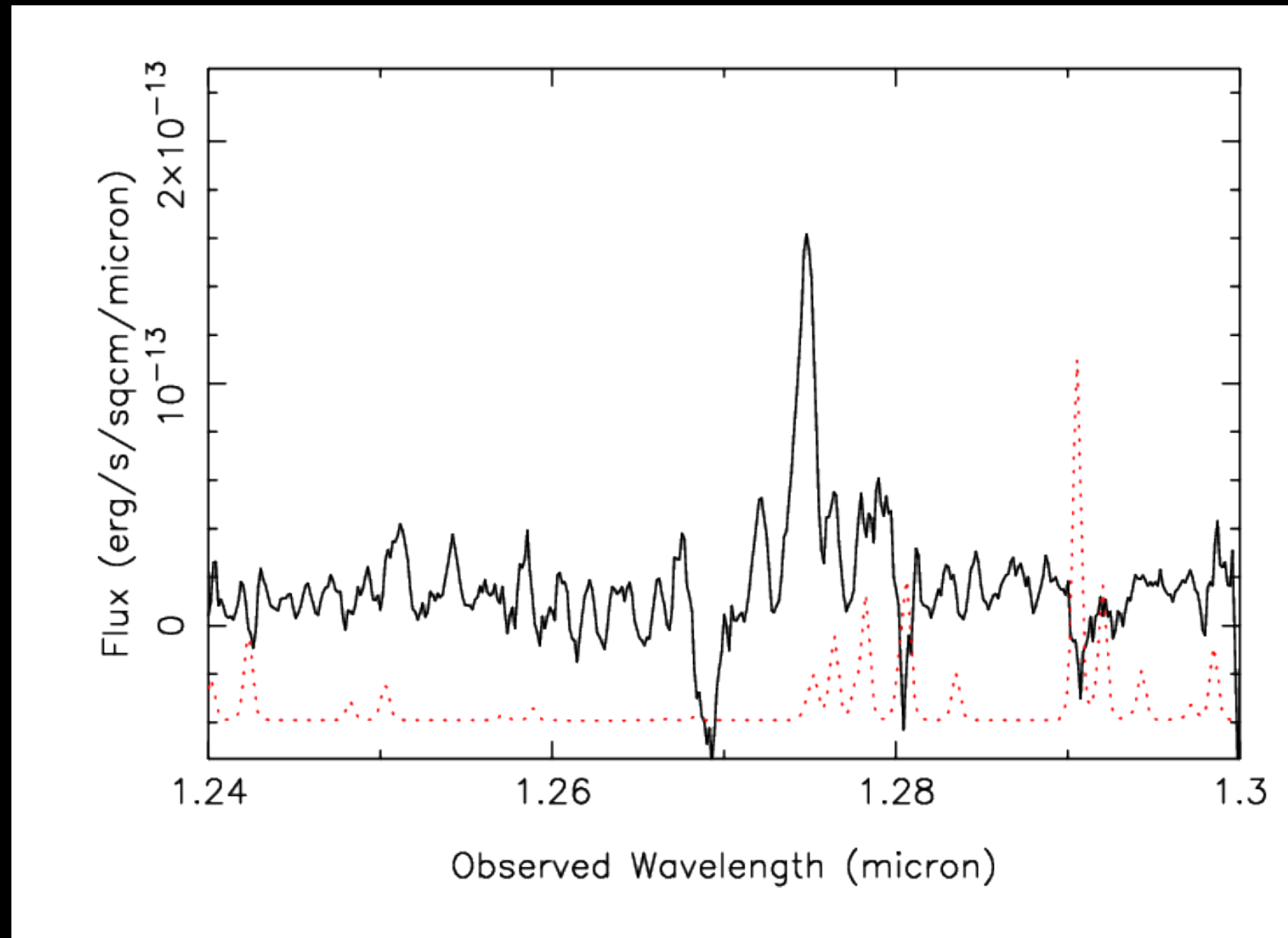
- $\text{SFR} \sim \text{few } M_{\odot}/\text{yr}$
- $b \sim 10\text{-}40 \text{ kpc}$

CP+ in prep



# Physical Properties

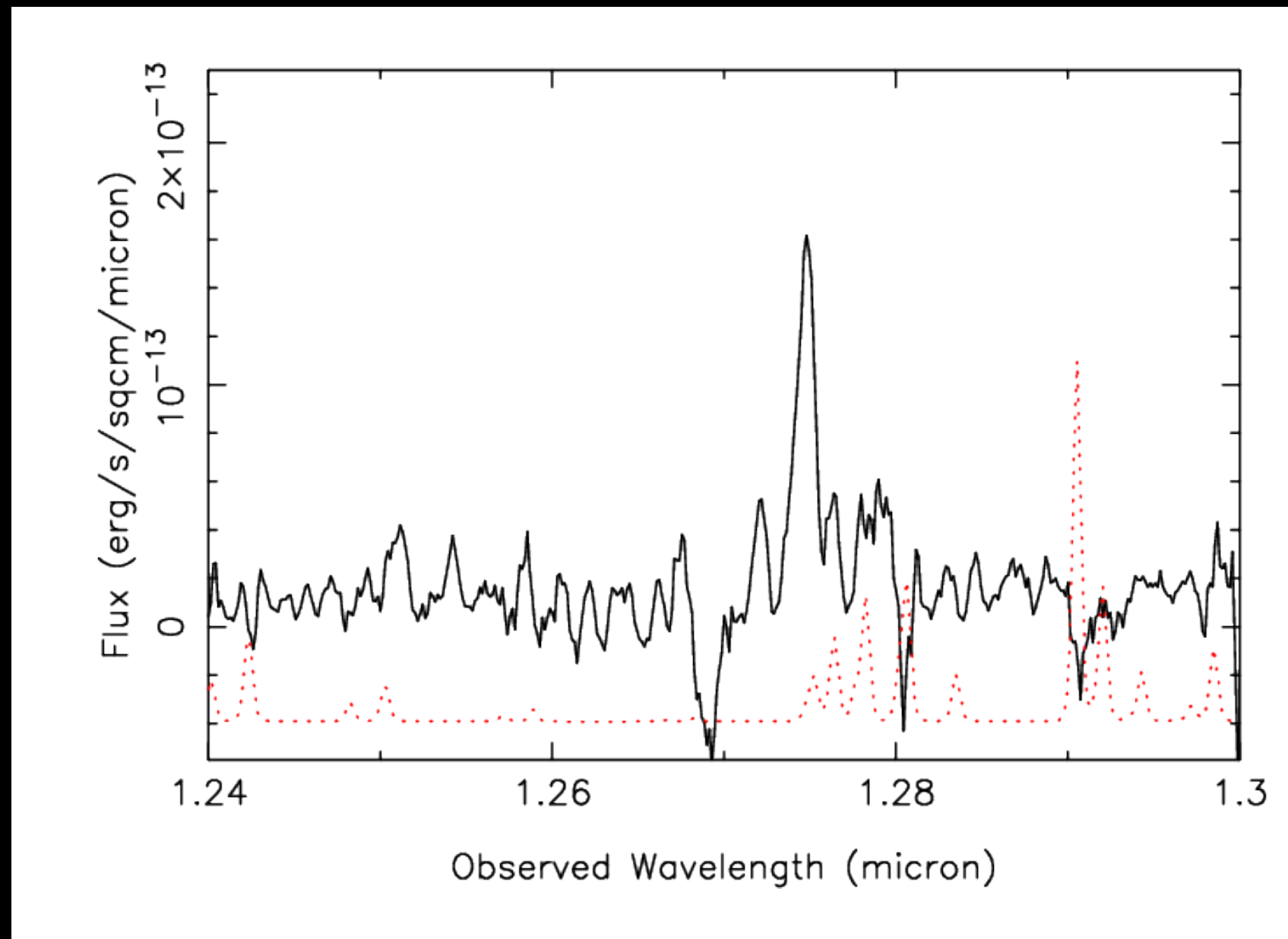
- $F(\text{H}\alpha) \sim 2 \times 10^{-16}$   
erg/s/cm<sup>2</sup>
- $\text{SFR} \sim 3.6 \pm 2.1$   
 $M_{\text{sun}}/\text{yr}$
- $b = 10.8''$   
 $b = 85 \text{ kpc}$



CP+ in prep

# Physical Properties

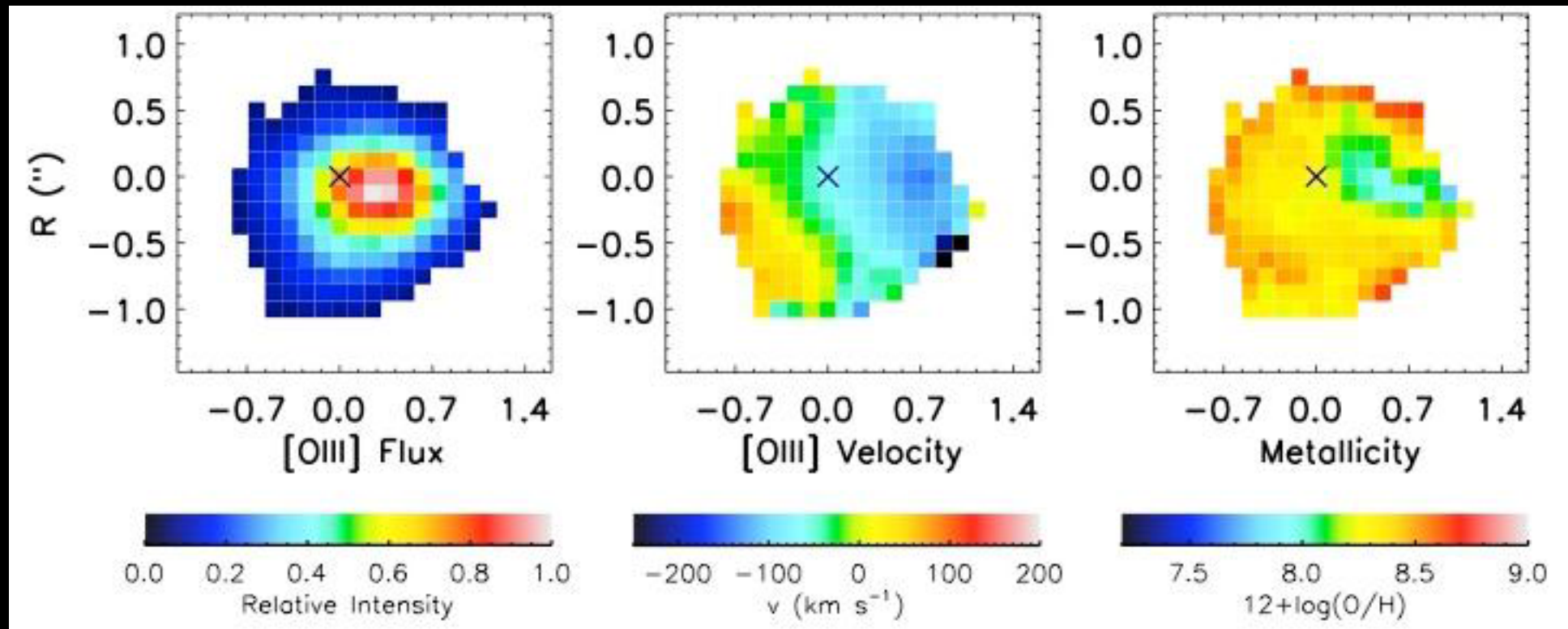
- N2 undetected
- [O/H] metallicity from N2 is  $12 + \log(\text{O}/\text{H}) < 9.20$



CP+ in prep



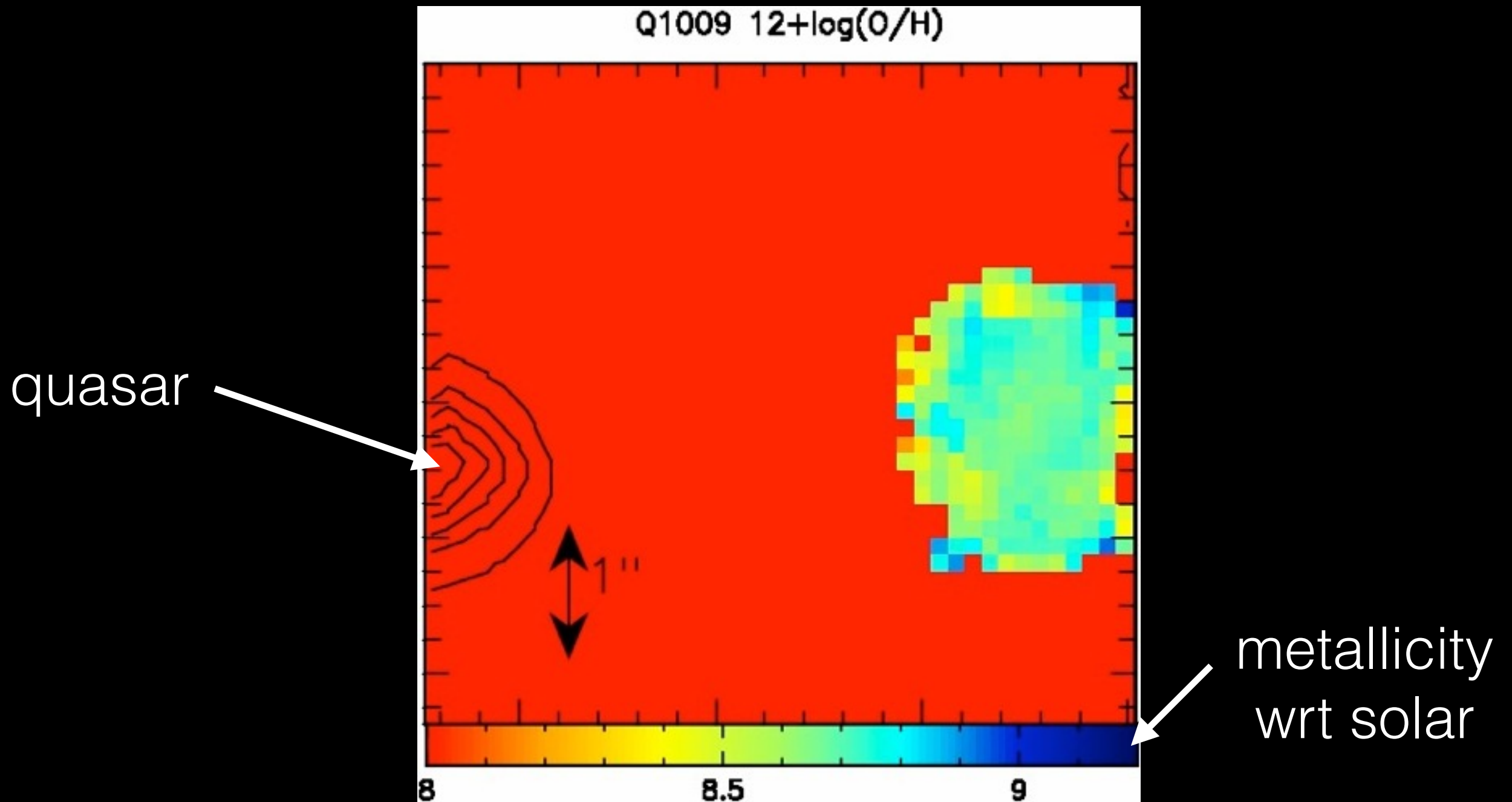
# Metallicity Gradients



- “inverted” metallicity gradients: accretion or merger?

Cresci+10, Nature, Epinat+12

# Metallicity Maps

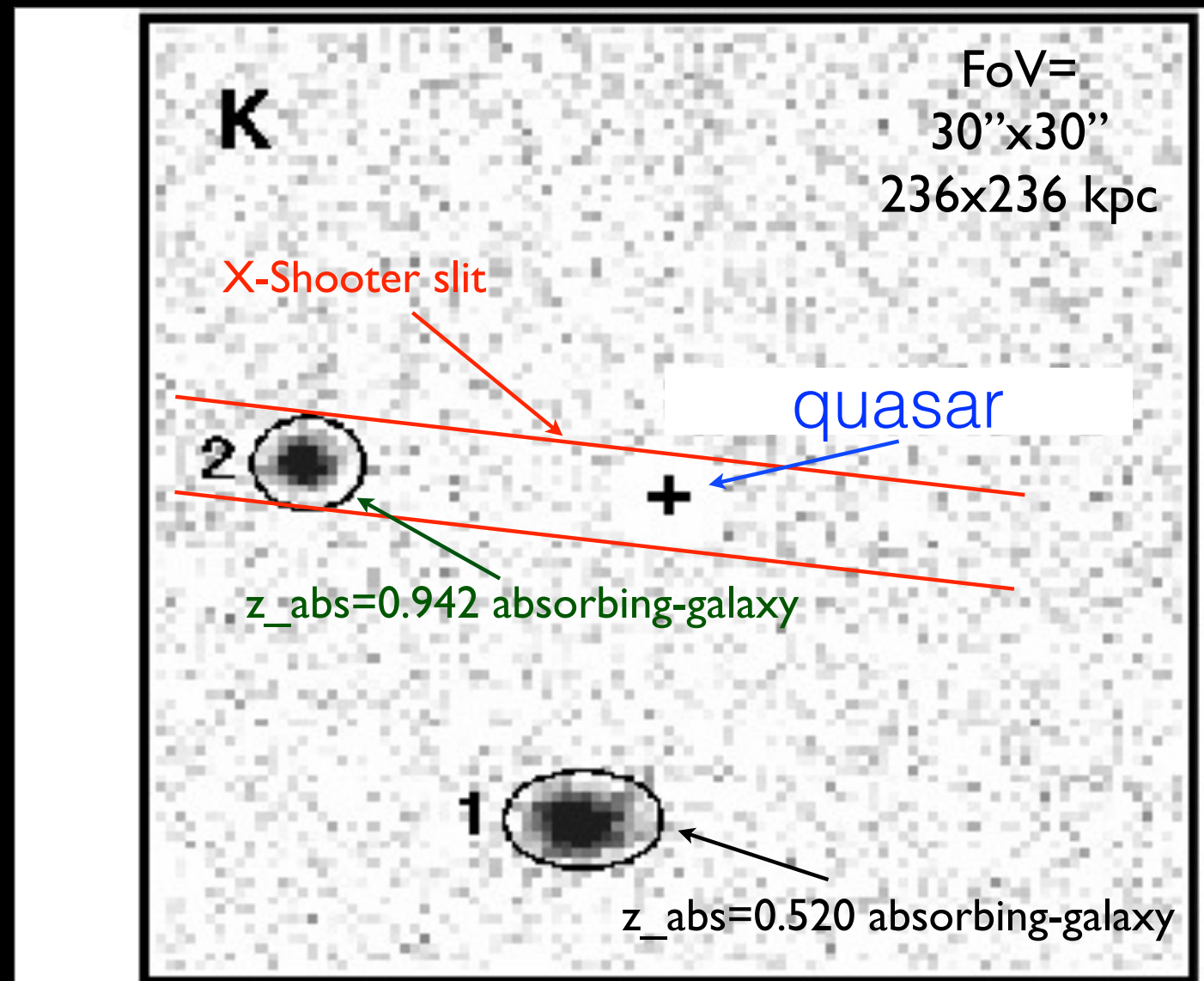


- gradients uniform

CP+11a

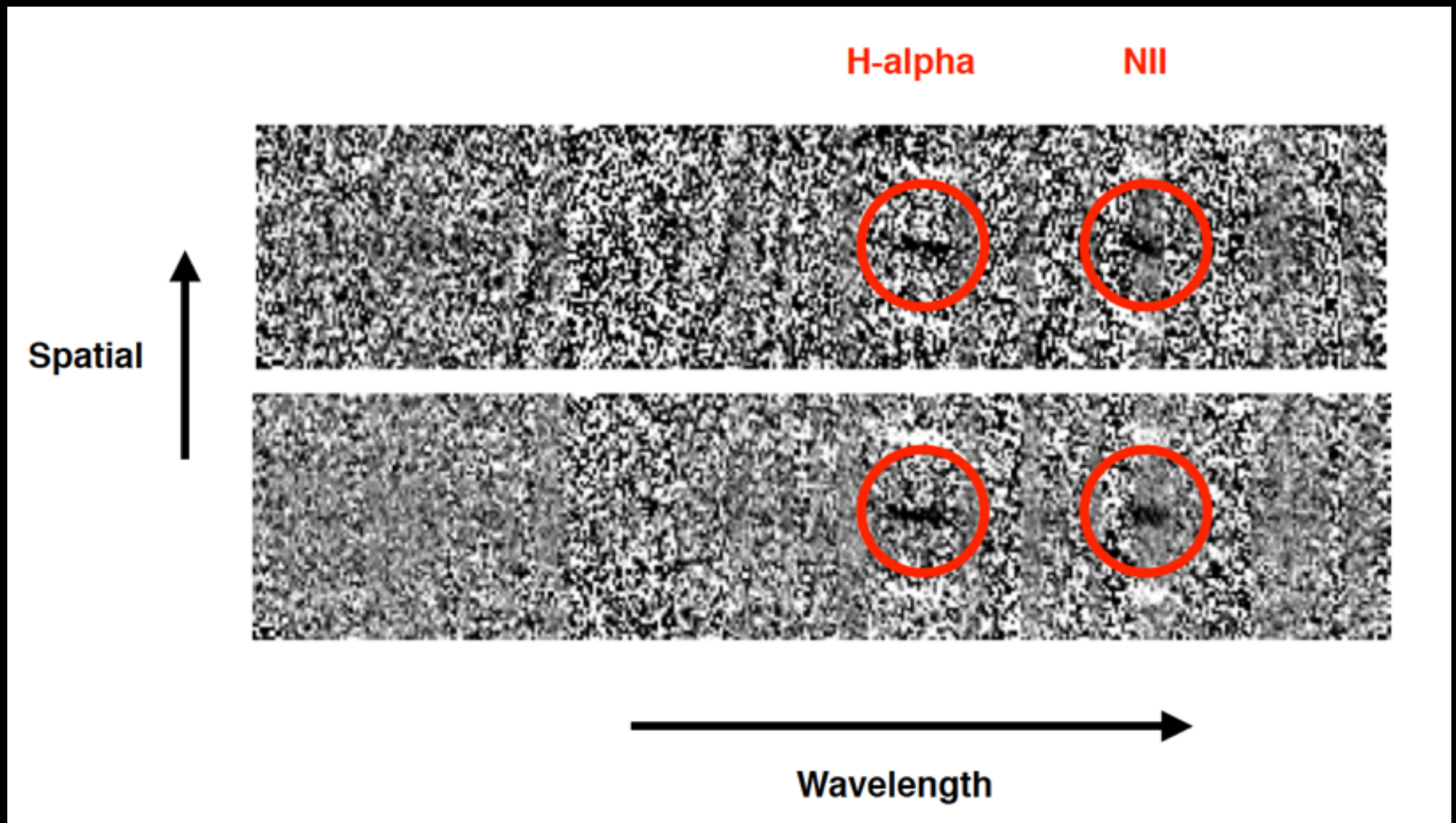
# Observational Strategy: Metallicity

- long slit covering both background quasar and foreground galaxy
- direct comparison of HI and HII gas



VLT/X-Shooter

# HII Emission Metallicity



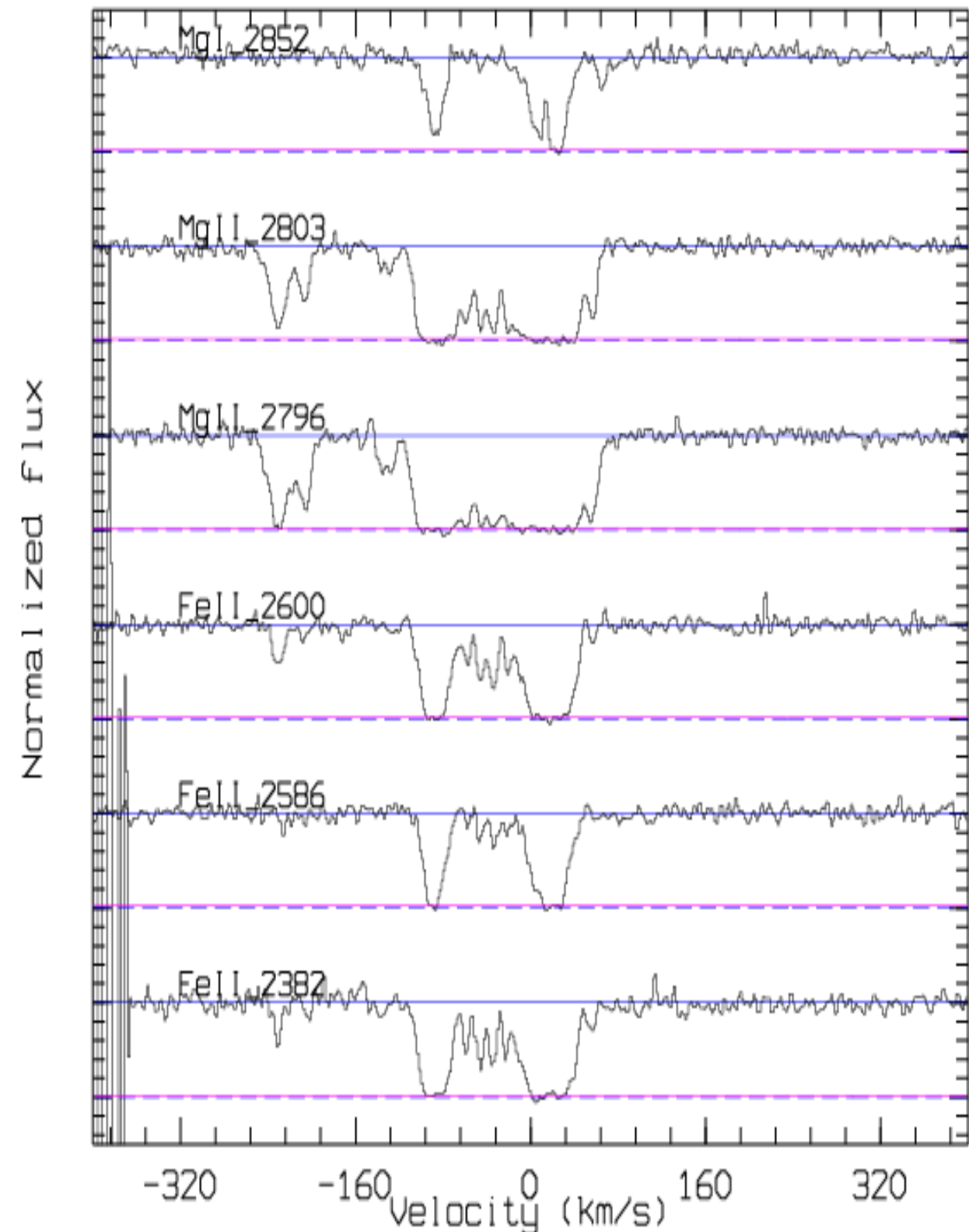
VLT/X-Shooter



# HI Neutral Gas Metallicity

- detecte AlIII, SiII, FeII, MgII and MgI  
Zn, Cr, Ti undetected
  - $[\text{Fe}/\text{H}] = -0.21 \pm 0.14$   
 $[\text{Zn}/\text{H}] < -0.41$
- Dessauges-Zavadsky+09

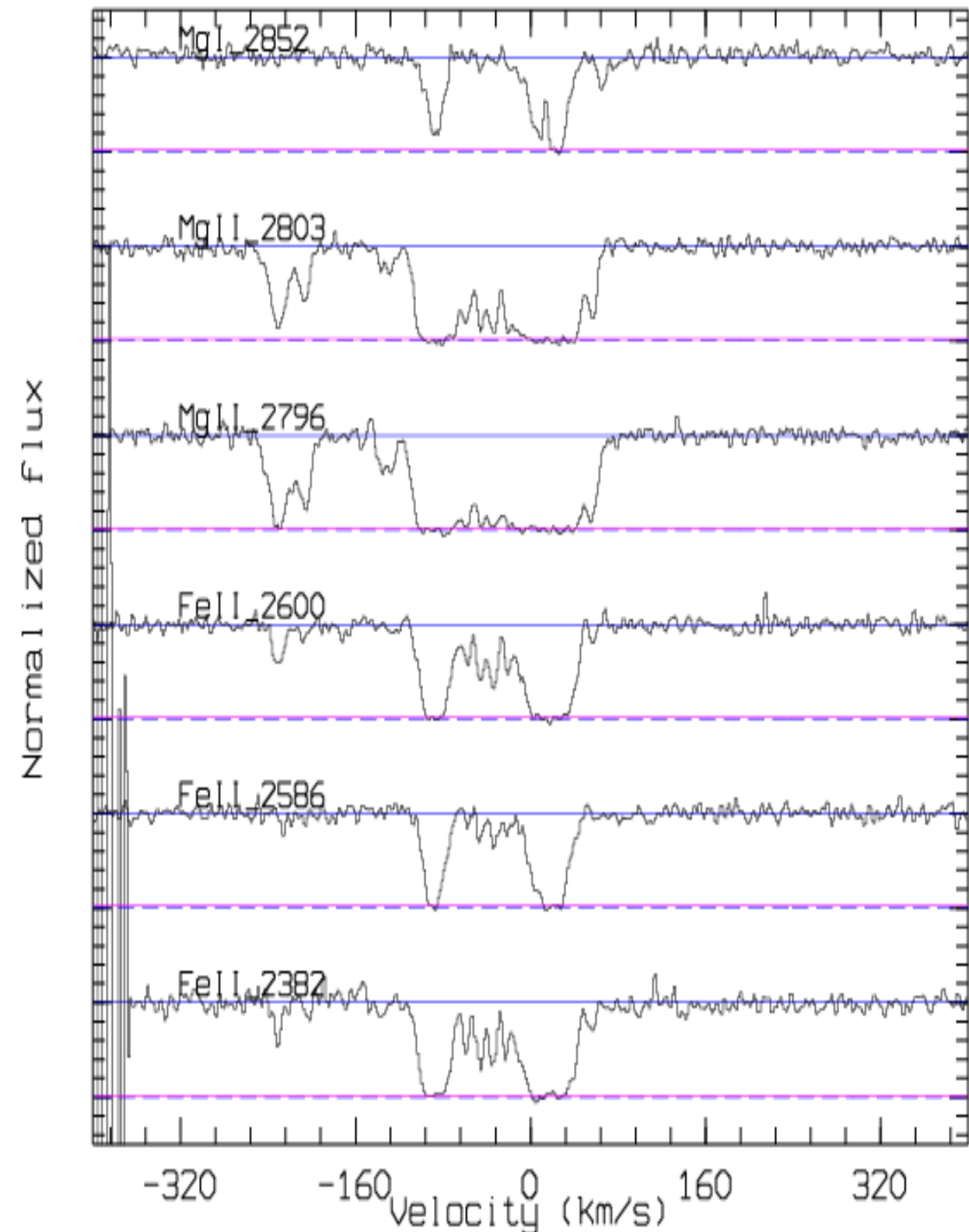
VLT/UVES



# HI Neutral Gas Kinematics

- $vel=0$  is  $z_{em}$
- double-peak profile
- additional blue-shifted component in the strongest absorption lines

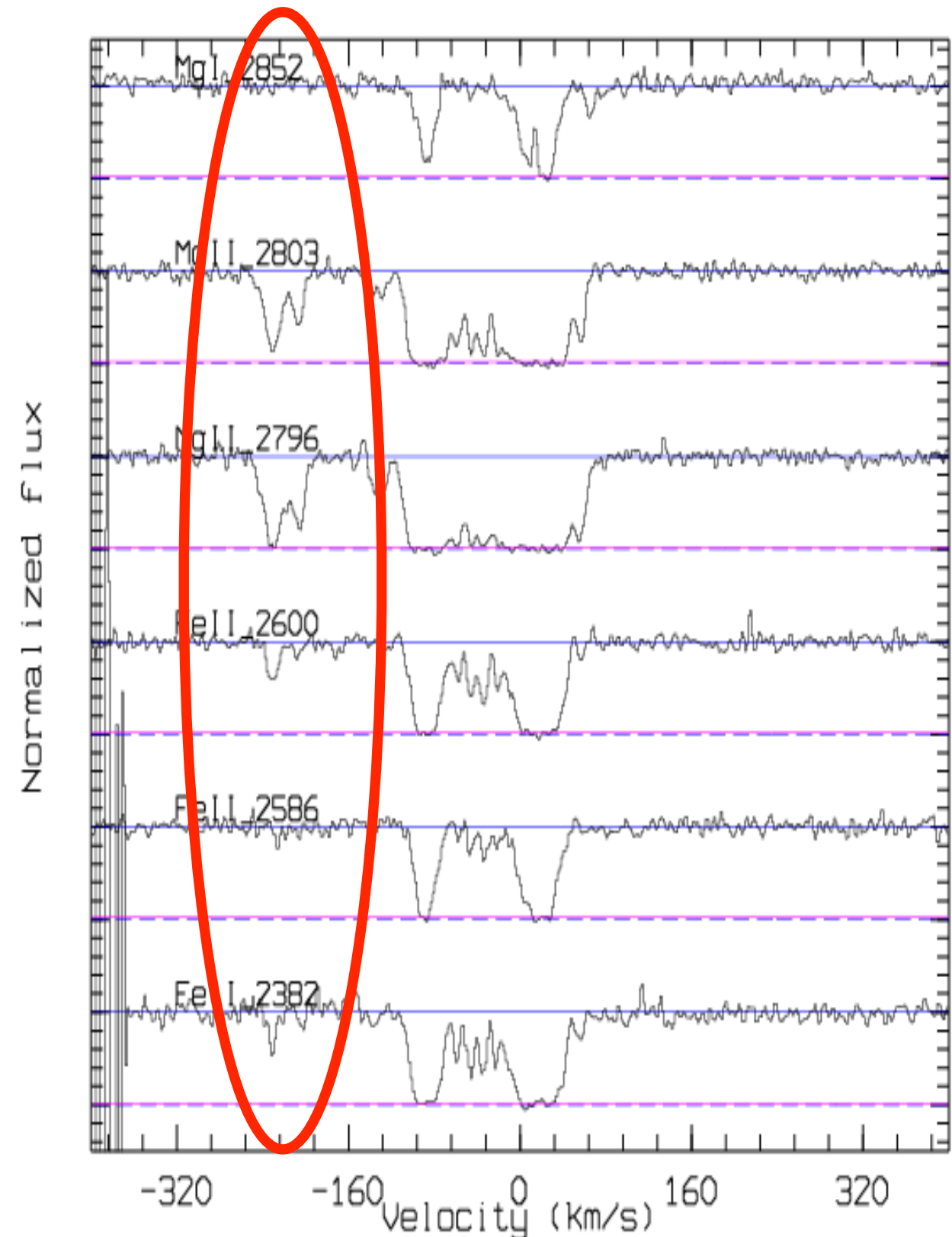
VLT/UVES



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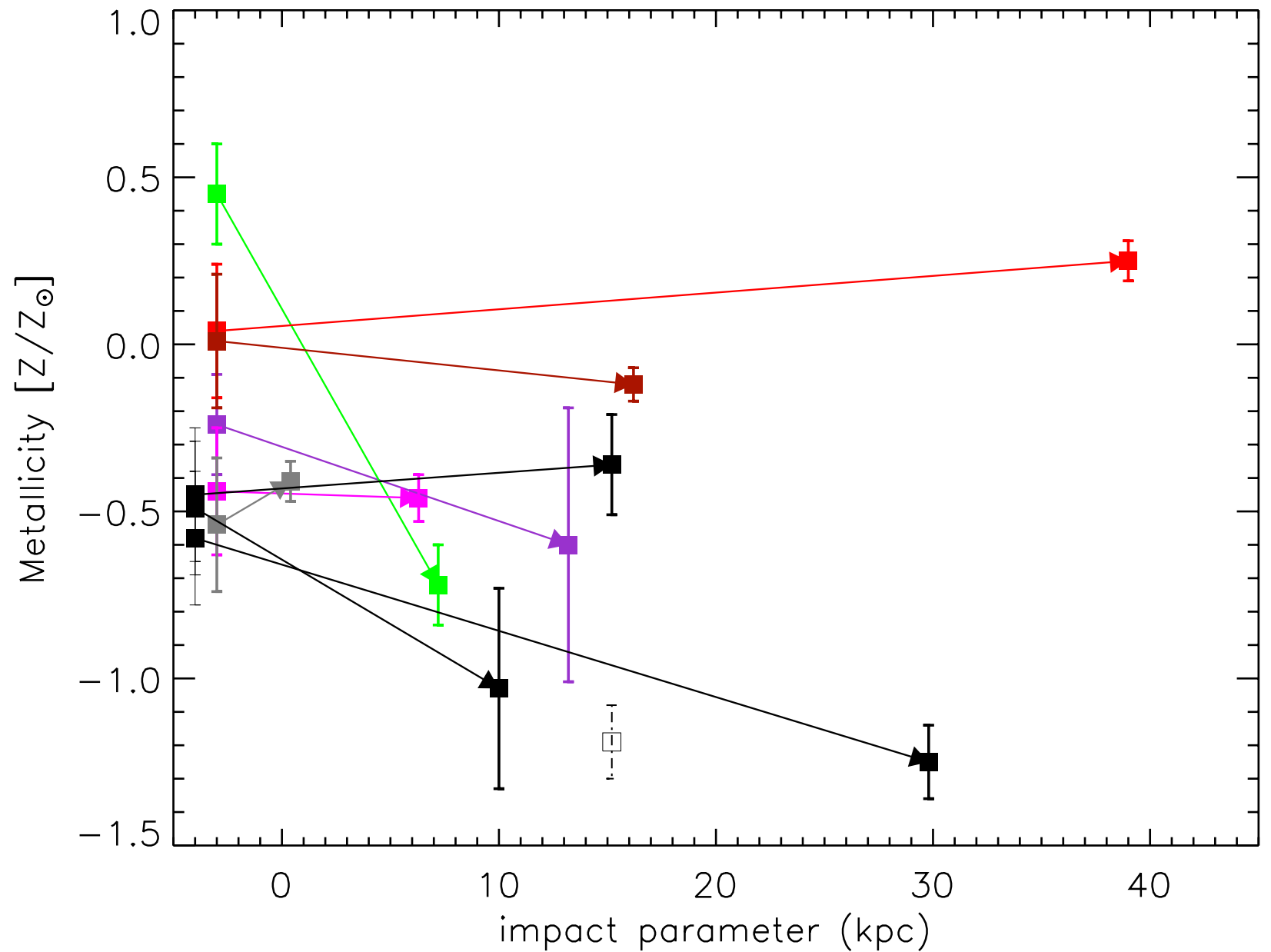




Hadi Rahmani

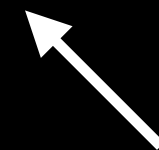
# Metallicity Gradients

HII emission  
metallicity



See Hadi's poster

$z \sim 0.6$



HI absorption  
metallicity



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# Conclusion

1. Neutral gas is constant across cosmic time:  
evidence for accretion on global scales
2. A combination of IFU and absorption techniques  
allow to characterise gas flows around galaxies