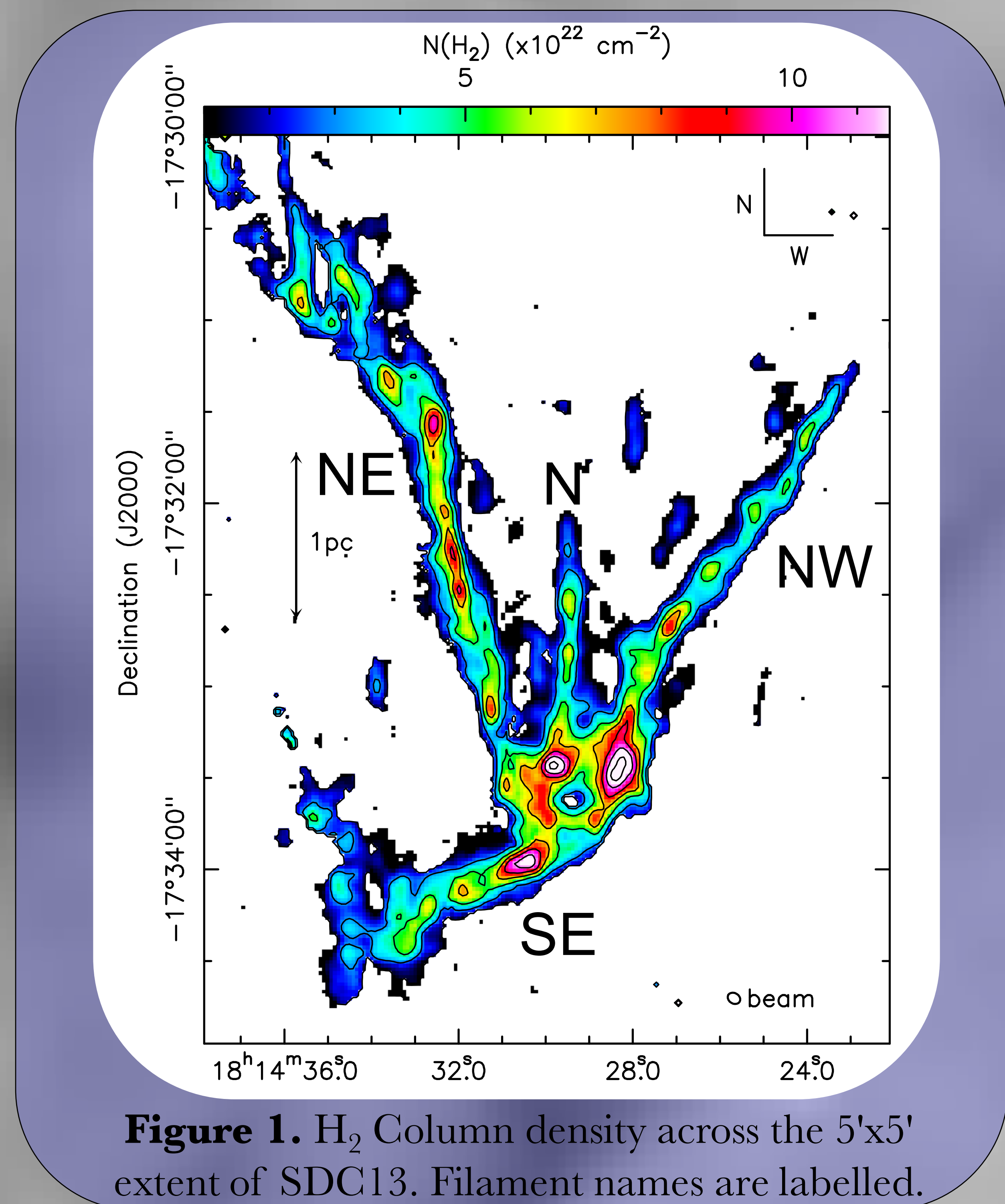


**Abstract:** We present new interferometer data of Ammonia molecular line emission towards SDC13, a unique interstellar filamentary cloud of molecular gas. For the first time, we observe gas velocity width peaks towards starless cores, opposite to the previously observed troughs in the “transition to coherence”. Also considering gravitational acceleration, altogether, these results hint at the importance of gravity in the evolution of infrared dark hub filament systems.

### 1. Introduction

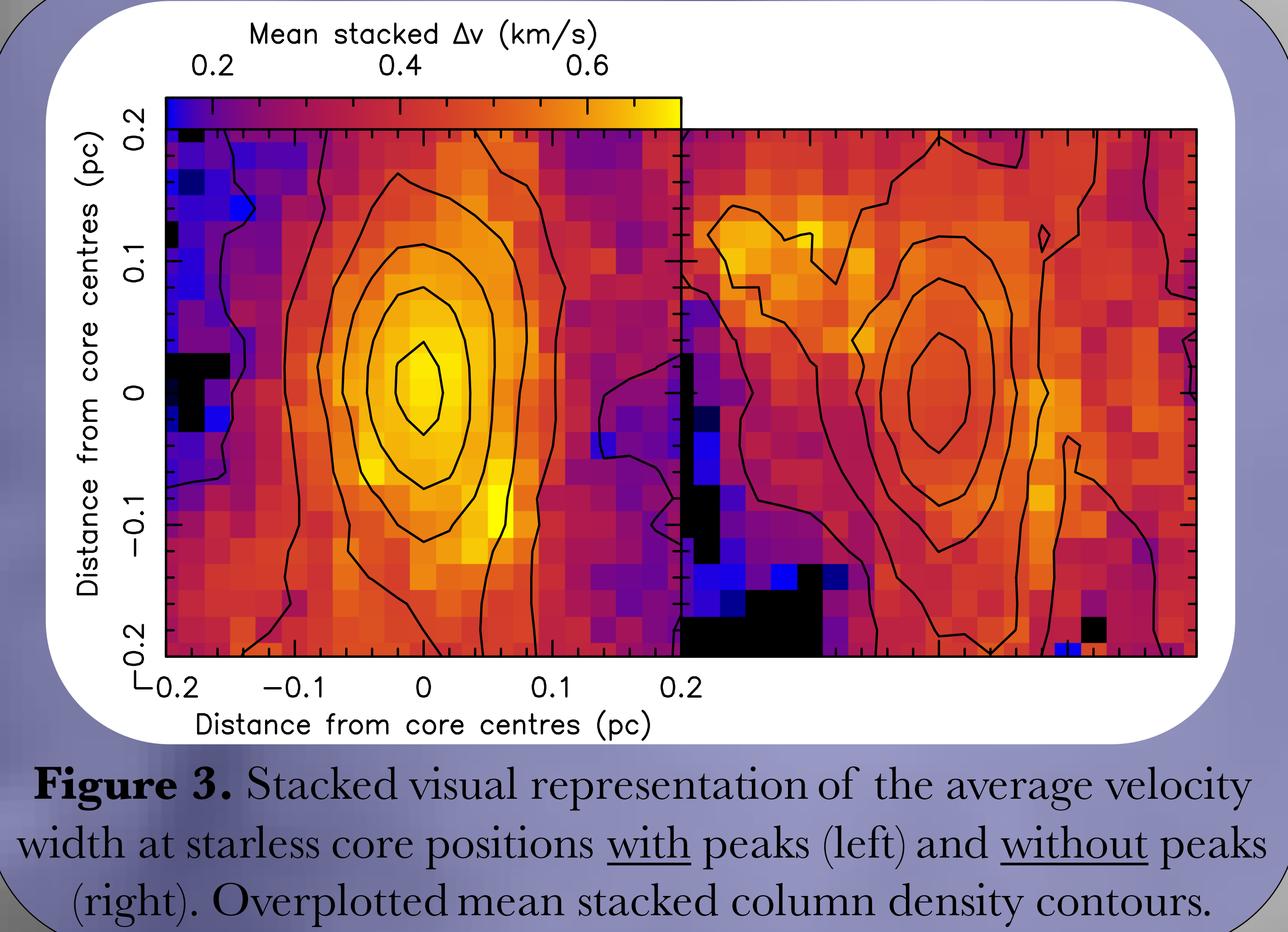
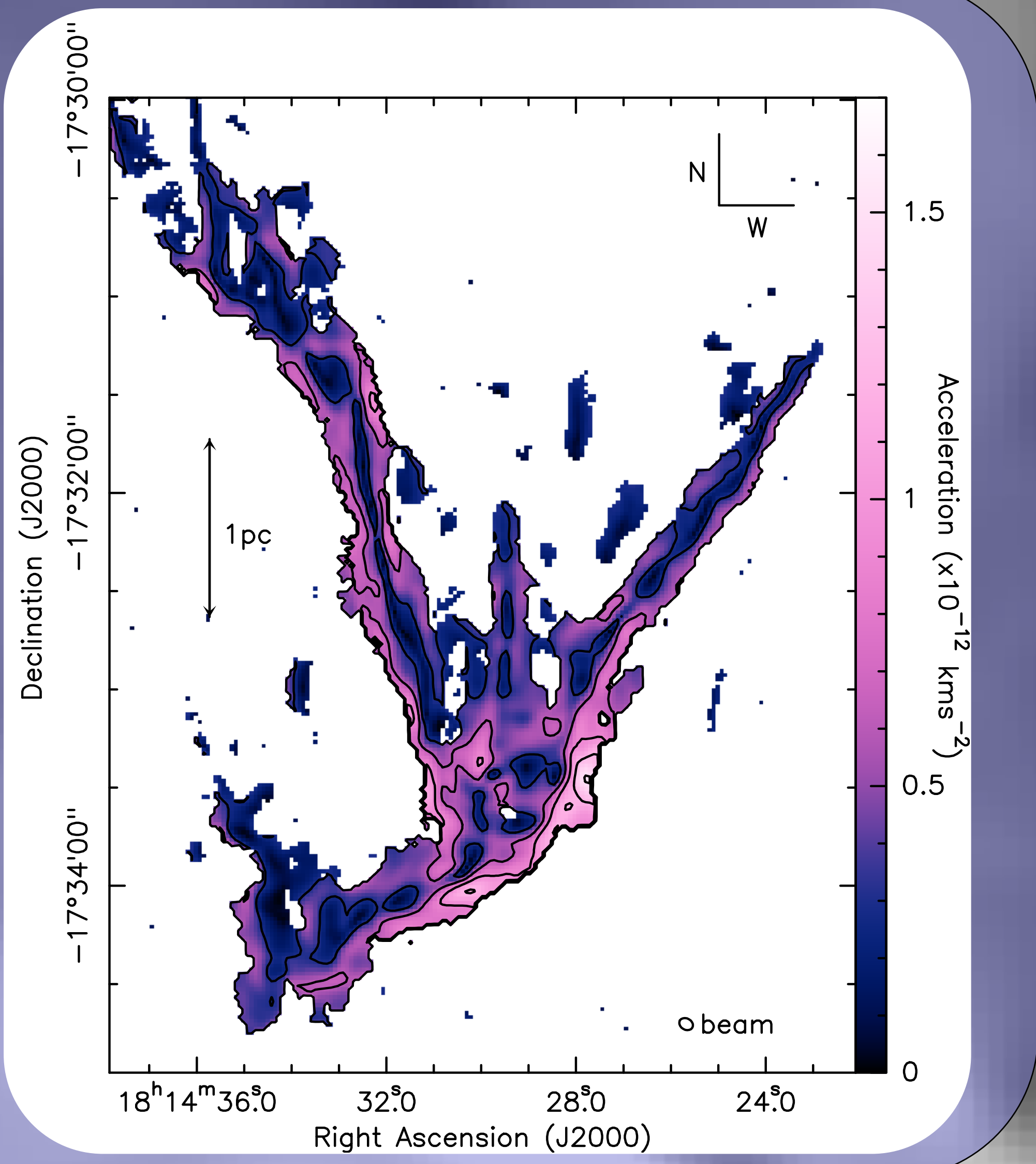
Interstellar filaments represent a key stage in star formation. As they become gravitationally unstable, the densest filaments fragment into cores [1], the direct progenitors of stars. The link between filament, core and star formation is one of the main issues of modern astronomy. Reservoirs of cold molecular gas called Infrared Dark Clouds (IRDCs) shed light on this as they contain the mostly pristine initial fingerprints of star formation.

SDC13 is a hub IRDC system of 4 parsec-long filaments located 3.6kpc away, containing  $1000M_{\odot}$  of material (Figure 1) [2] [3]. We present new interferometer data of SDC13 down to spatial scales of  $\sim 0.07\text{pc}$ , capable of probing the missing link between filament, core and star formation [4].

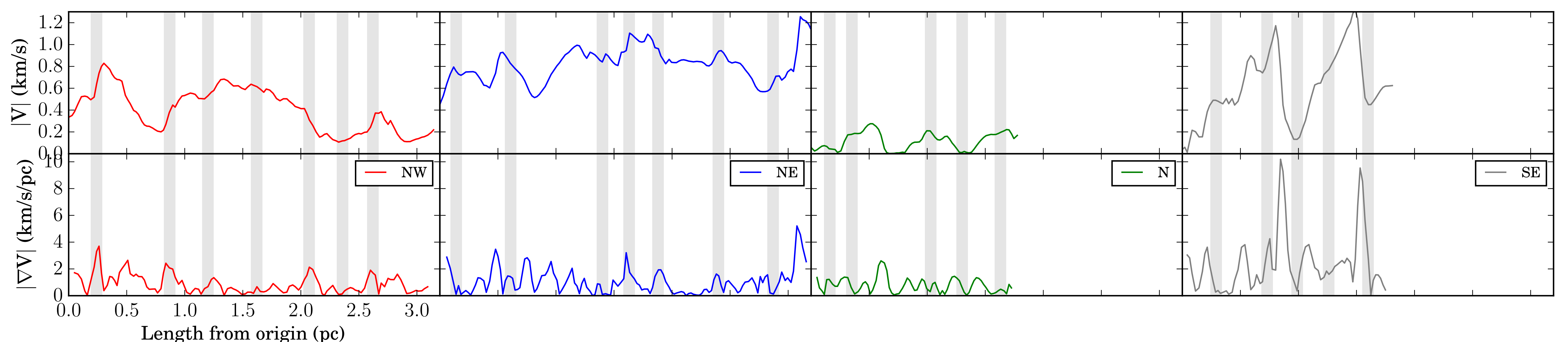


**Figure 1.**  $\text{H}_2$  Column density across the  $5' \times 5'$  extent of SDC13. Filament names are labelled.

**Figure 2.** The acceleration due to gravity at each pixel, where the mass of each pixel was calculated from the  $\text{H}_2$  column density of Figure 1.



**Figure 3.** Stacked visual representation of the average velocity width at starless core positions with peaks (left) and without peaks (right). Overplotted mean stacked column density contours.



**Figure 4.** Profiles along the spine lengths of filaments: NW (red), NE (blue), N (green) and SE (grey). The top panel shows the  $\text{H}_2$  column density, the middle shows the absolute line-of-sight velocity, while the bottom shows the absolute velocity gradient. Vertical shaded regions represent core positions.

### 2. Results

Our main results are:

- Regular  $\sim 0.33\text{pc}$  core spacing
- Velocity width peaks at 61% of starless cores
- Starless core GPE to KE conversion efficiency  $\epsilon \sim 20\%$
- Velocity gradient peaks at  $\sim 50\%$  of cores, as expected for accreting cores
- Strongest acceleration gradient towards position of largest core at the hub centre

**References** [1] André et al. 2010, AAP, 518, L102. [2] Peretto & Fuller 2009, AAP, 505, 405. [3] Peretto et al. 2014, A&A, 561, A83. [4] Williams et al. 2017, in preparation. Background: SDC13 Spitzer  $8\mu\text{m}$  map [3].

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